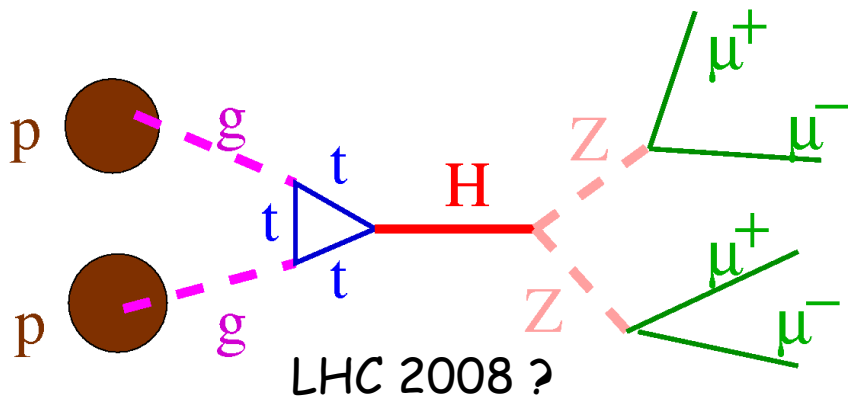
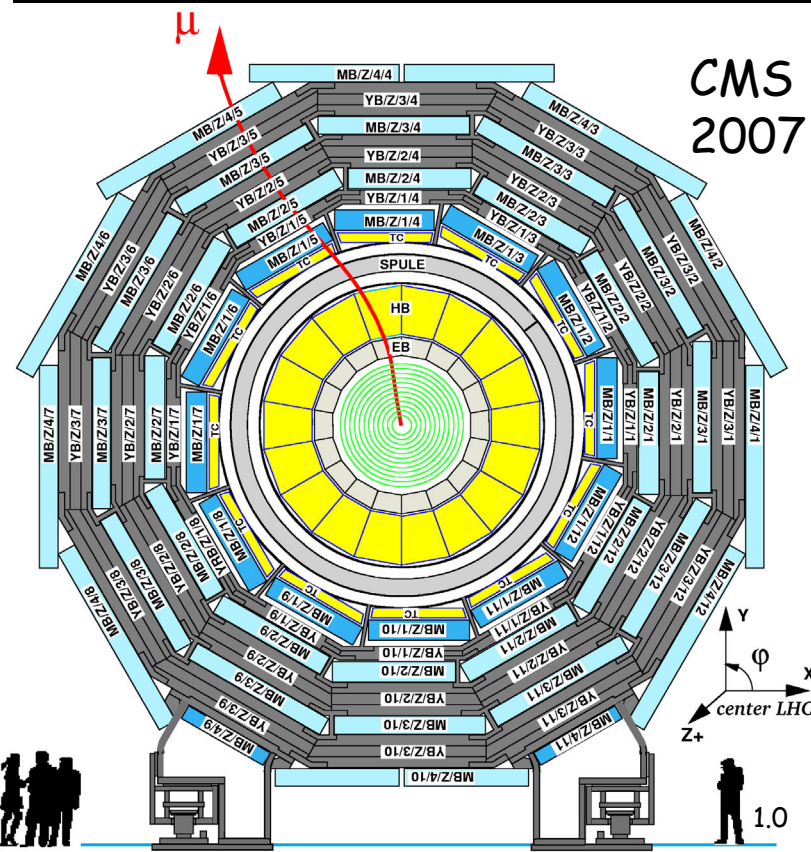
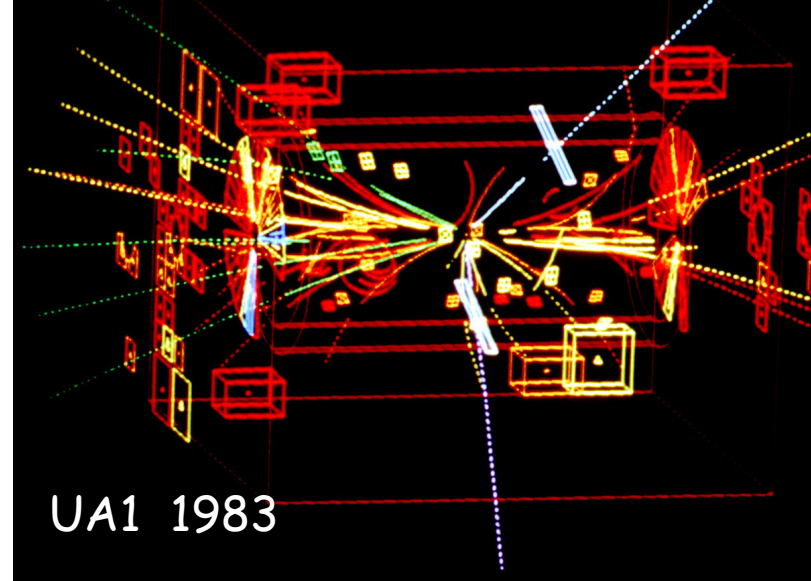


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Part I Introduction

Part II Standard Model Physics

Part III Higgs

Part IV New Phenomena

- SUSY
 - motivation
 - searches:
 - R-Parity conserved
 - R-Parity violated
- Extra dimensions
- Black holes

References

SUperSymmetry

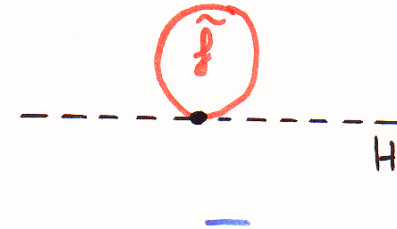
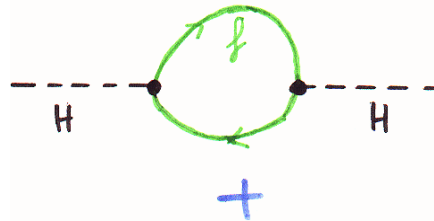
Particle	Spin	Susy-Partner	Spin
ν_e	1/2	$\tilde{\nu}_e^L$	0
e^-	1/2	$\tilde{e}_L^-, \tilde{e}_R^-$	0
u	1/2	\tilde{u}_L, \tilde{u}_R	0
d	1/2	\tilde{d}_L, \tilde{d}_R	0
γ, Z, h, H, A	1, 0	$\tilde{\chi}_1^0, \tilde{\chi}_2^0, \tilde{\chi}_3^0, \tilde{\chi}_4^0$	1/2
W^\pm, H^\pm	1, 0	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^\pm$	1/2
g	1	\tilde{g}	1/2

if R-parity (-1 for sparticles) is conserved:
 \Rightarrow LSP = Lightest SUSY particle = $\tilde{\chi}_1$ = stable

SUSY

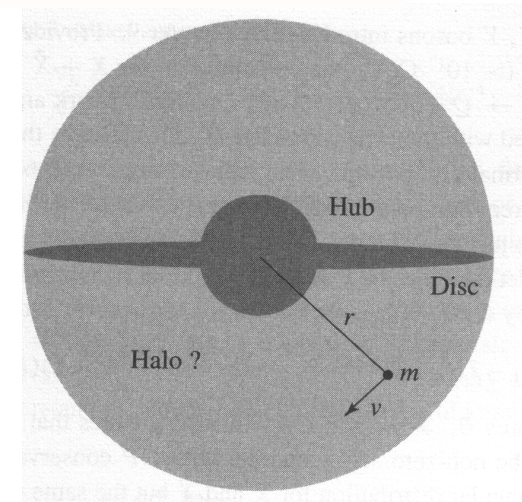
Nice features:

- symmetry relating **bosons (forces)** \leftrightarrow fermions (matter)
- higgs mass m_H
under control
- grand unification (incl. gravity!) possible
- neutralino = dark matter candidate



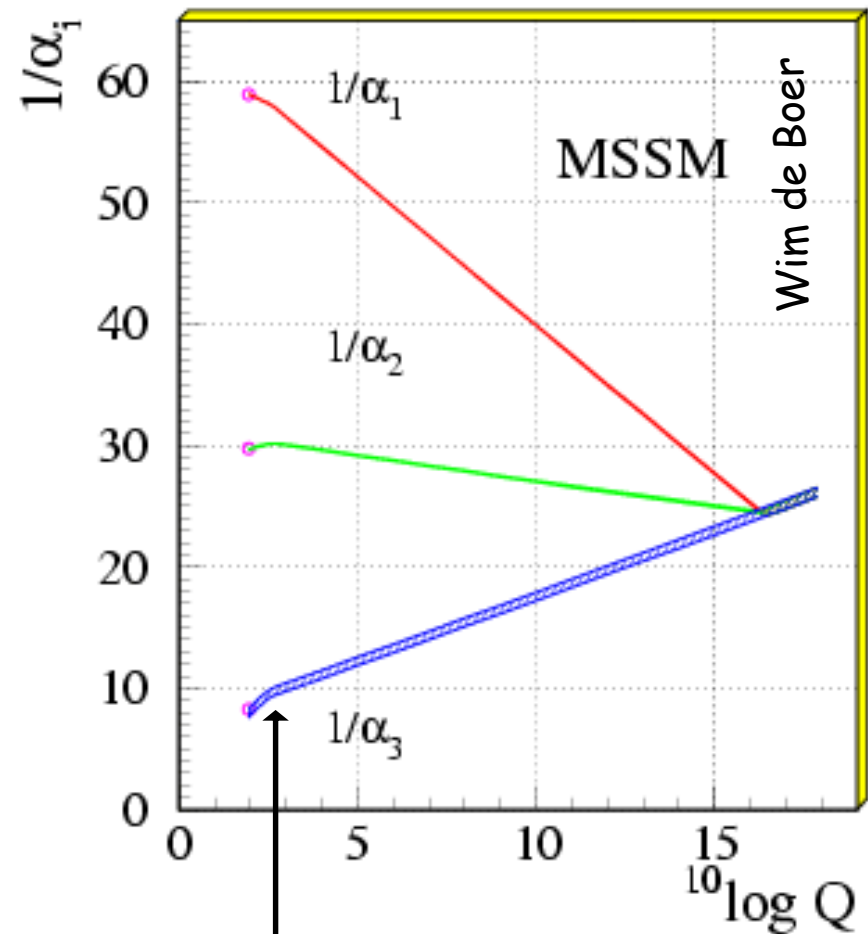
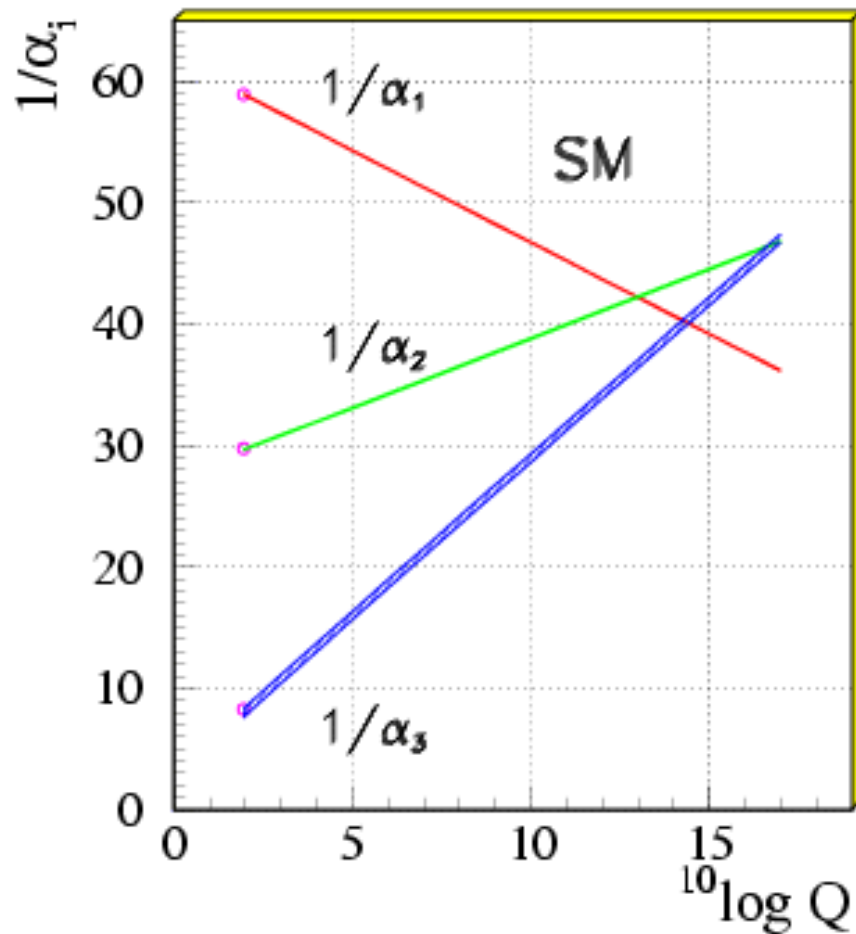
Other properties:

- SUSY broken: **no sparticle seen yet**
- > 100 new parameters \Rightarrow **Minimal Model (MSSM, MSUGRA)**



Grand Unification ?

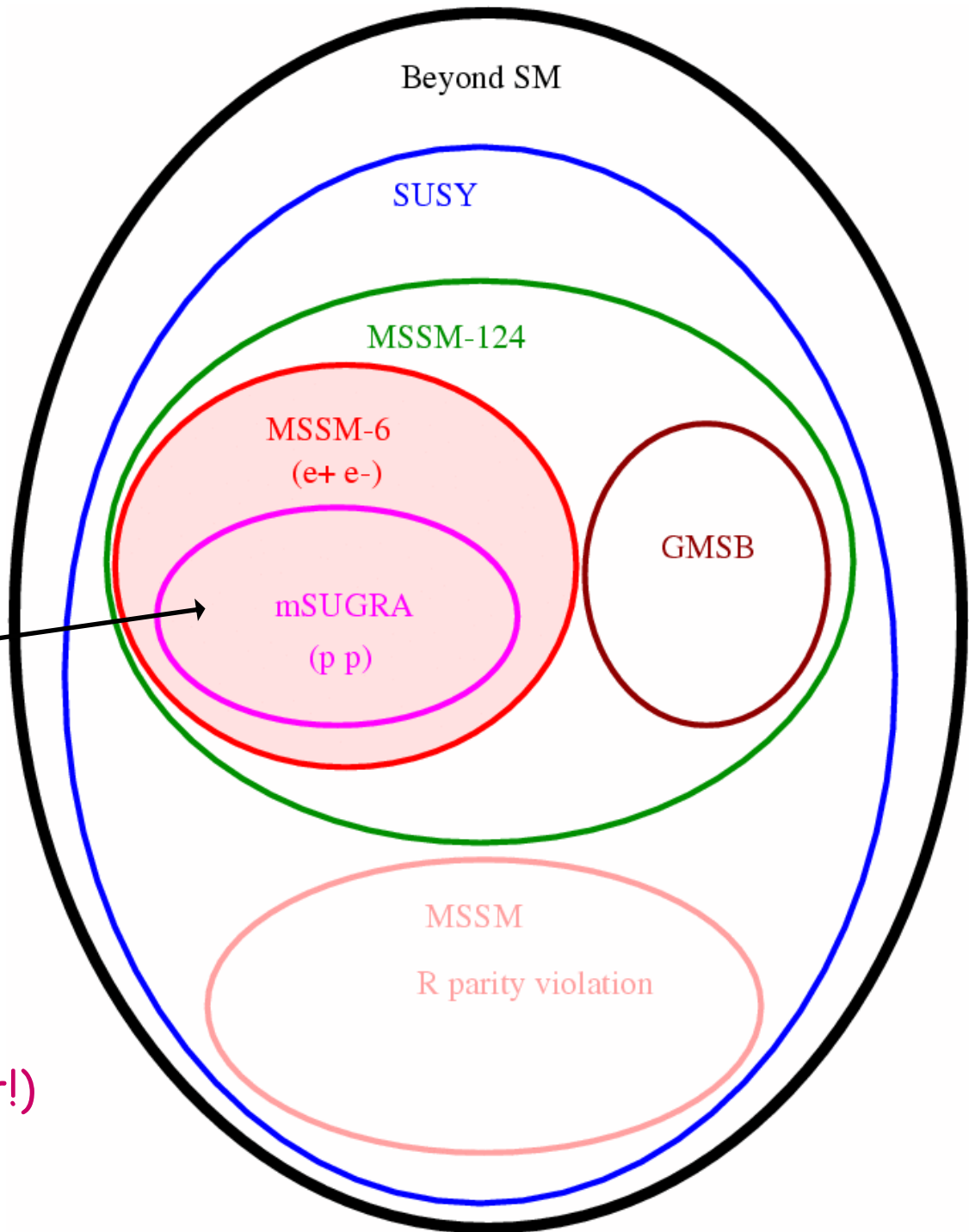
Unification of the Coupling Constants in the SM and the minimal MSSM



SUSY mass scale \sim TeV

Wim de Boer

SUSY models



investigated
most often
by pp experiments

(does not mean it's right!)

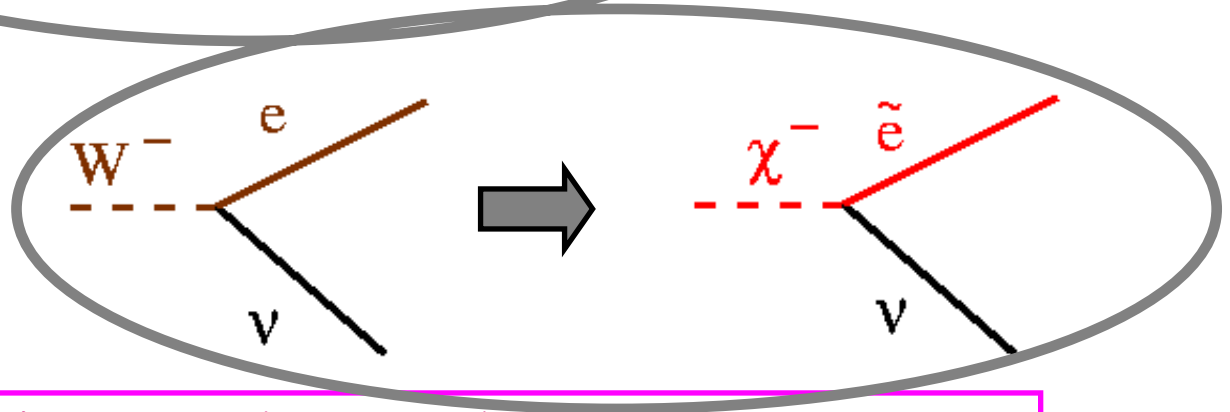
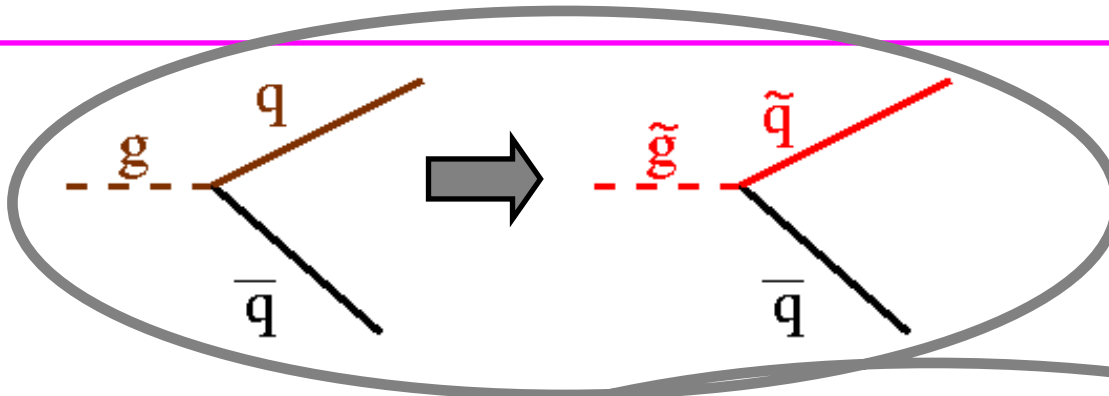
SUSY interactions (R conserved)

Feynman graphs:

take any SM vertex with 3 or 4 particles

replace two legs by the corresponding sparticles (\sim)

Examples



Coupling constants (electroweak, strong):

same as in SM !

MSUGRA parameters

MSUGRA = Minimal SuperGRAvity model

m_0 = universal scalar mass at GUT scale (s..., higgs)

$m_{1/2}$ = universal gaugino mass at GUT scale (...inos)

$\tan \beta$ = v_2/v_1 = ratio of higgs vacuum expectation values

A_0 = universal sfermion mass mixing parameter [GUT]

$\text{sgn } \mu$ = sign of higgsino mass parameter

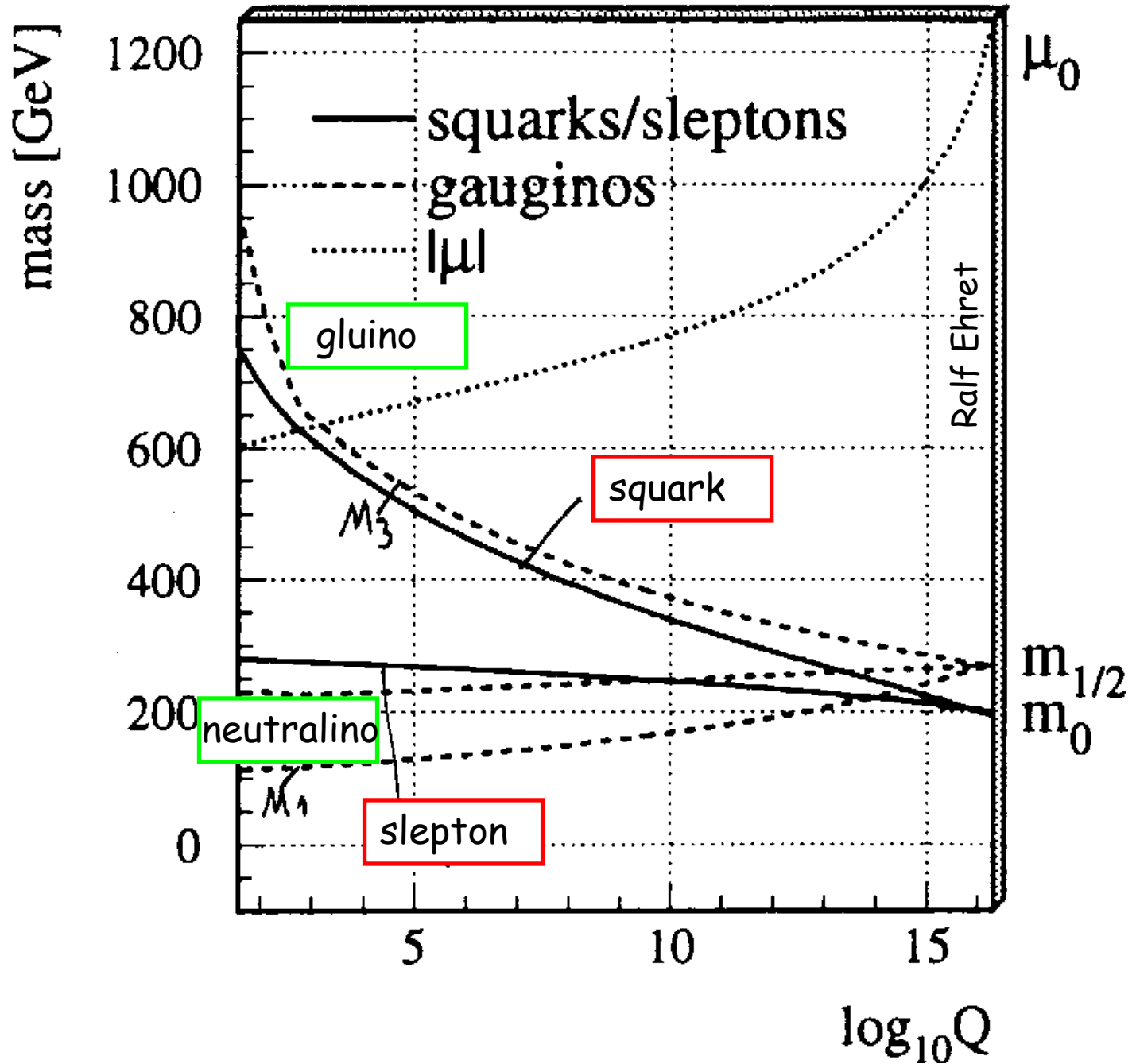
Required:

$M(\text{SUSY}) < 1 \text{ TeV}$

LSP without electromagnetic and strong coupling

Note: m_h given by m_0 ... LEP higgs limit = severe constraint

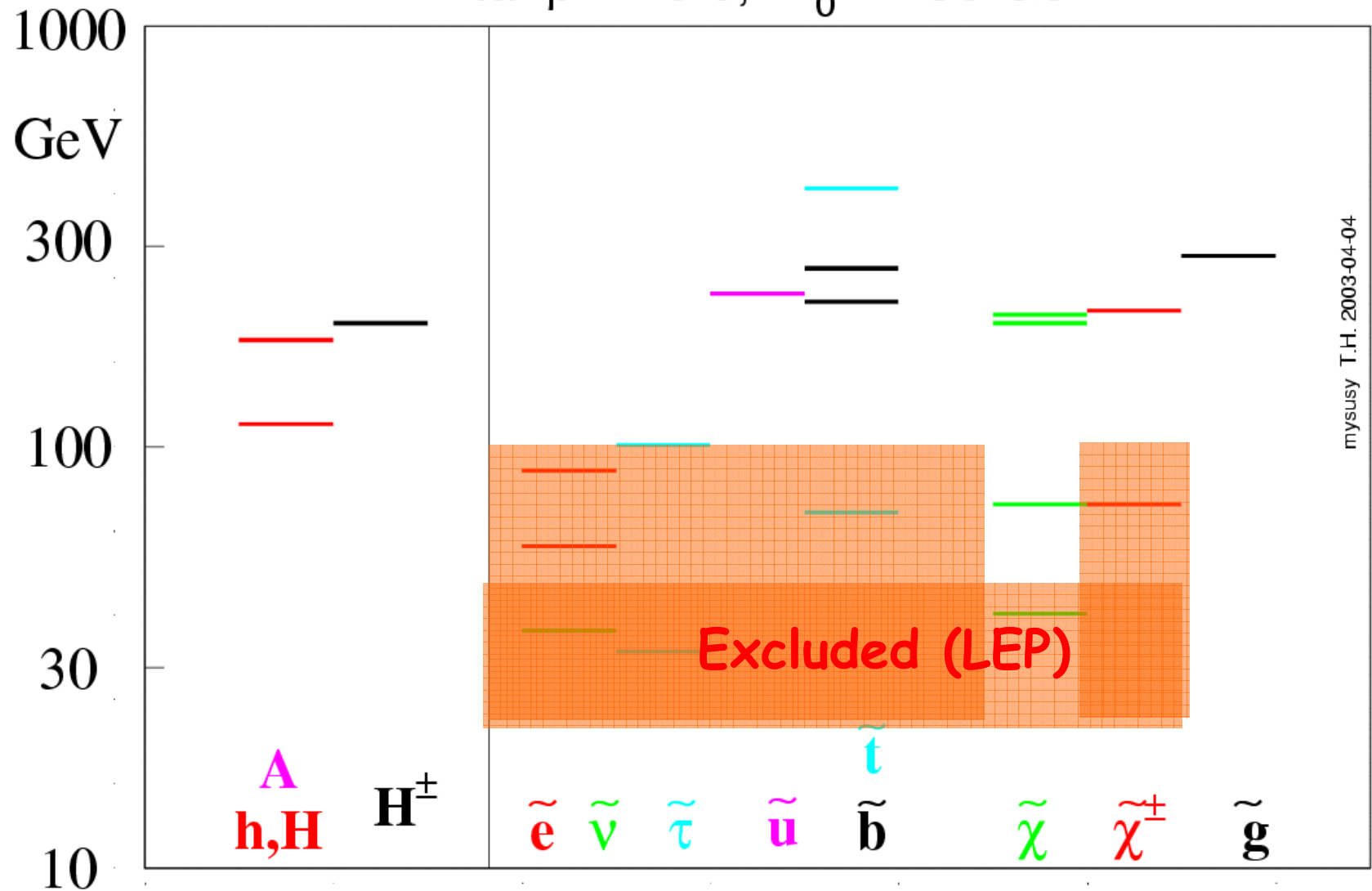
MSUGRA masses



MSUGRA scenario 1

$$m_0 = 10 \text{ GeV}, \quad m_{1/2} = 100 \text{ GeV}, \quad \mu \text{ neg.}$$

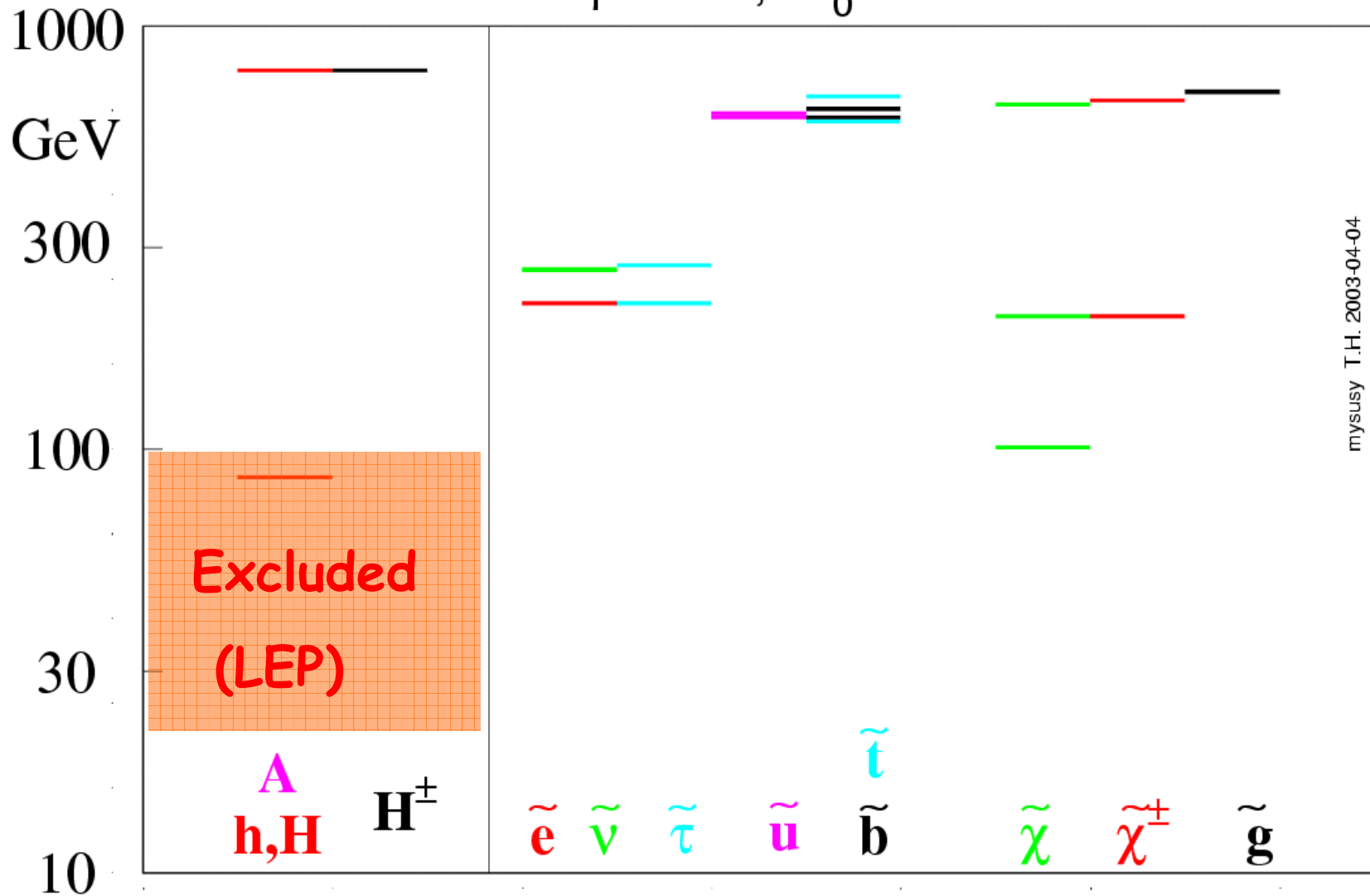
$$\tan\beta = 10.0, \quad A_0 = 450 \text{ GeV}$$

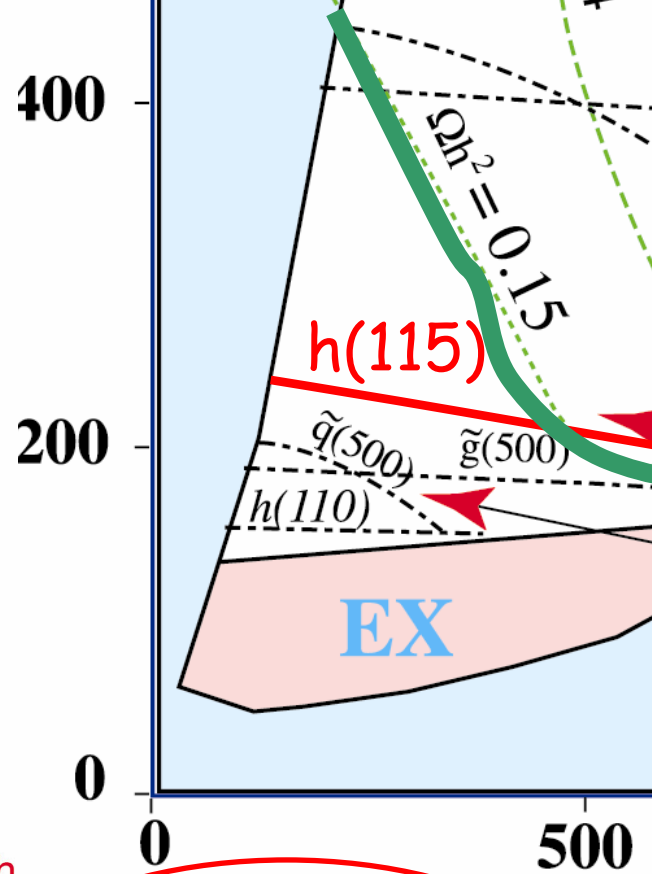
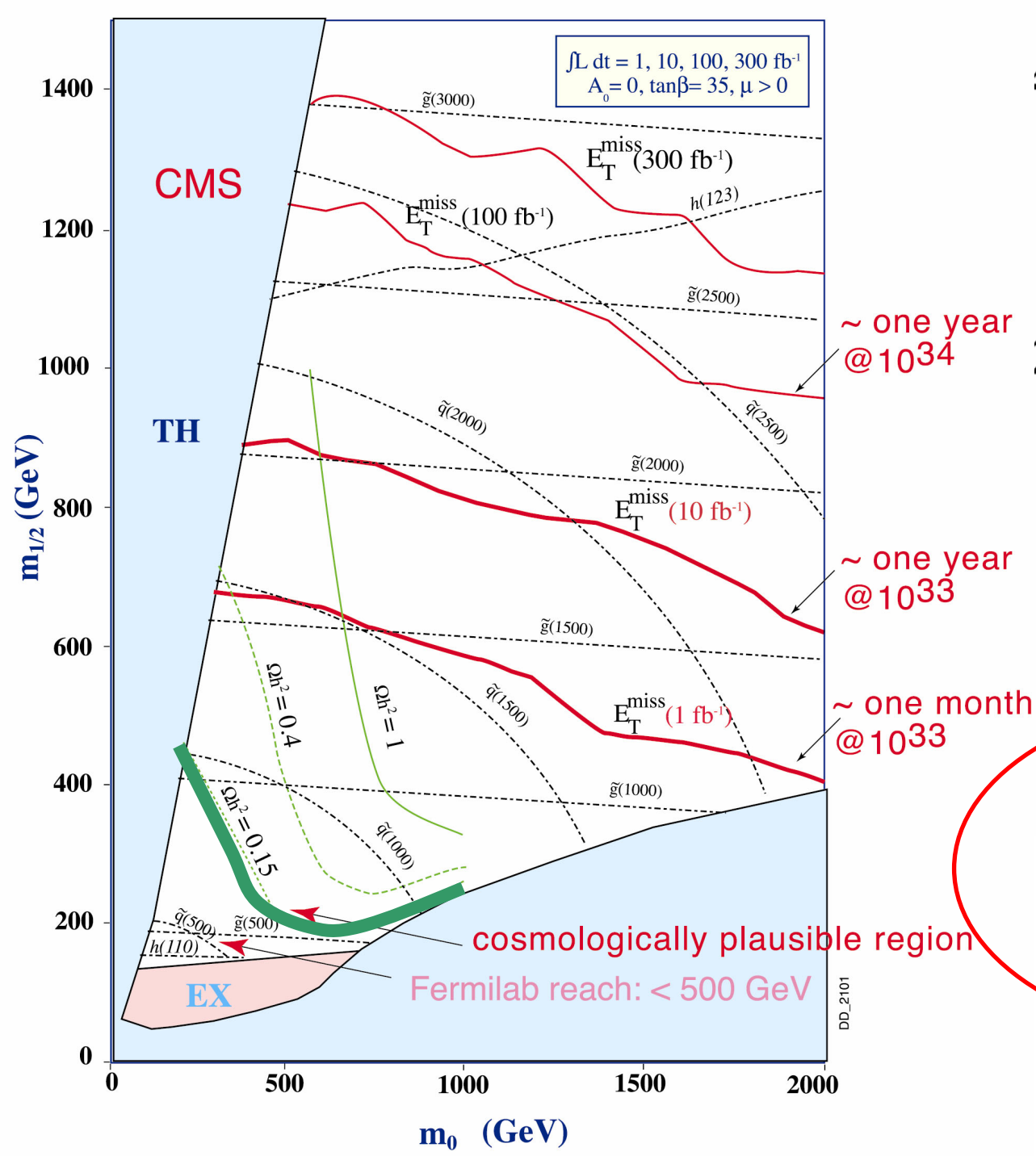


MSUGRA scenario 2

$$m_0 = 200 \text{ GeV}, \quad m_{1/2} = 243 \text{ GeV}, \quad \mu \text{ neg.}$$

$$\tan\beta = 2.0, \quad A_0 = 0 \text{ GeV}$$





Cosmological
 constraint (dark
 matter)
 essential!

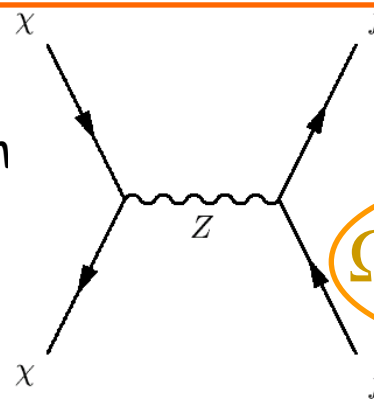
SUSY reach

Cosmological Constraints I

Assume: neutralino = $\tilde{\chi}_0$ = dark matter = WIMP

Early universe:

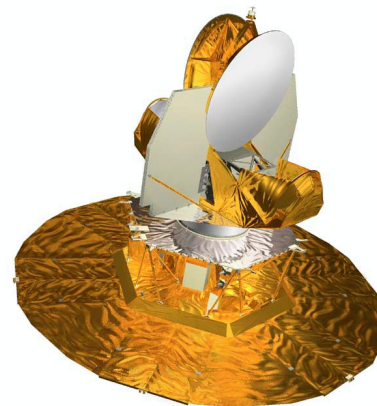
- theory
- 1) Very high temperature and pressure:
creation and annihilation: equilibrium
 - 2) High temperature and pressure:
annihilation dominates: $N(\tilde{\chi}_0) \downarrow$
 - 3) Low temperature and pressure:
freeze out: $N(\tilde{\chi}_0) = \text{const}$



$$\Omega_{DM} = f(m_\chi, \dots)$$

observation

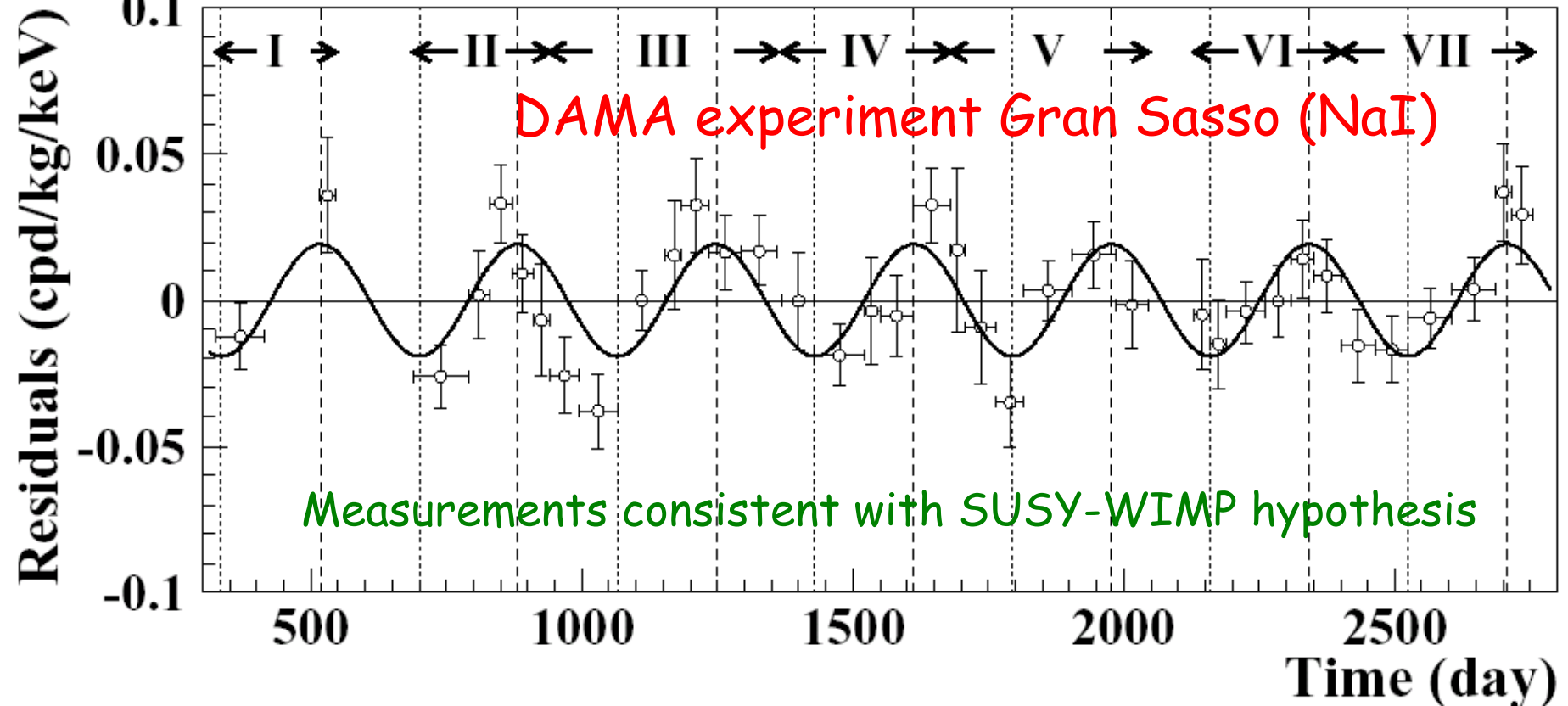
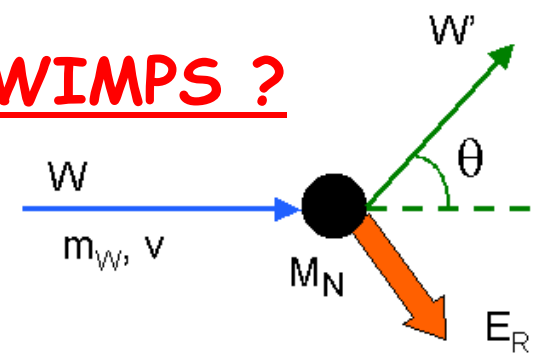
WMAP
measurement of
cosmic microwave
background

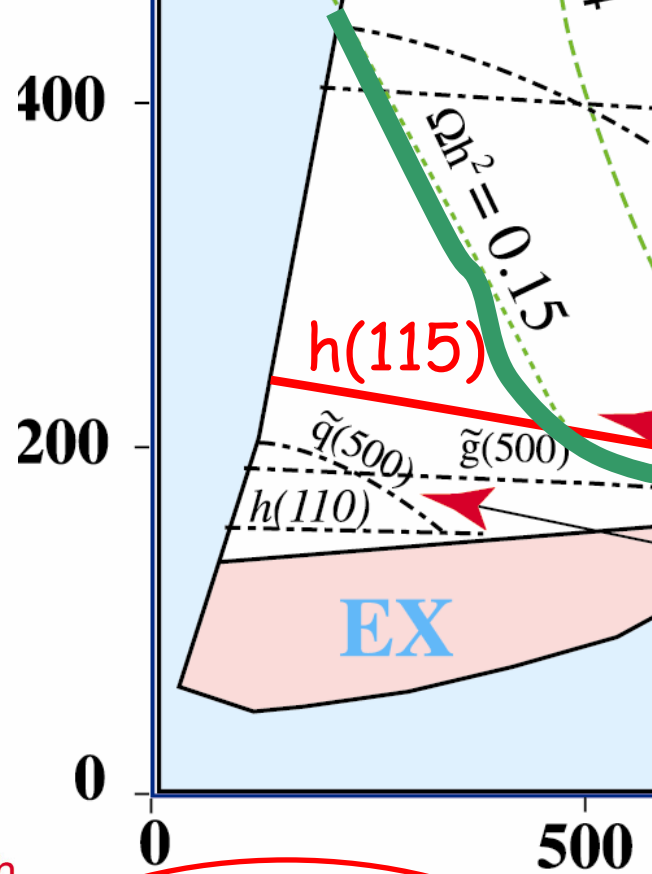
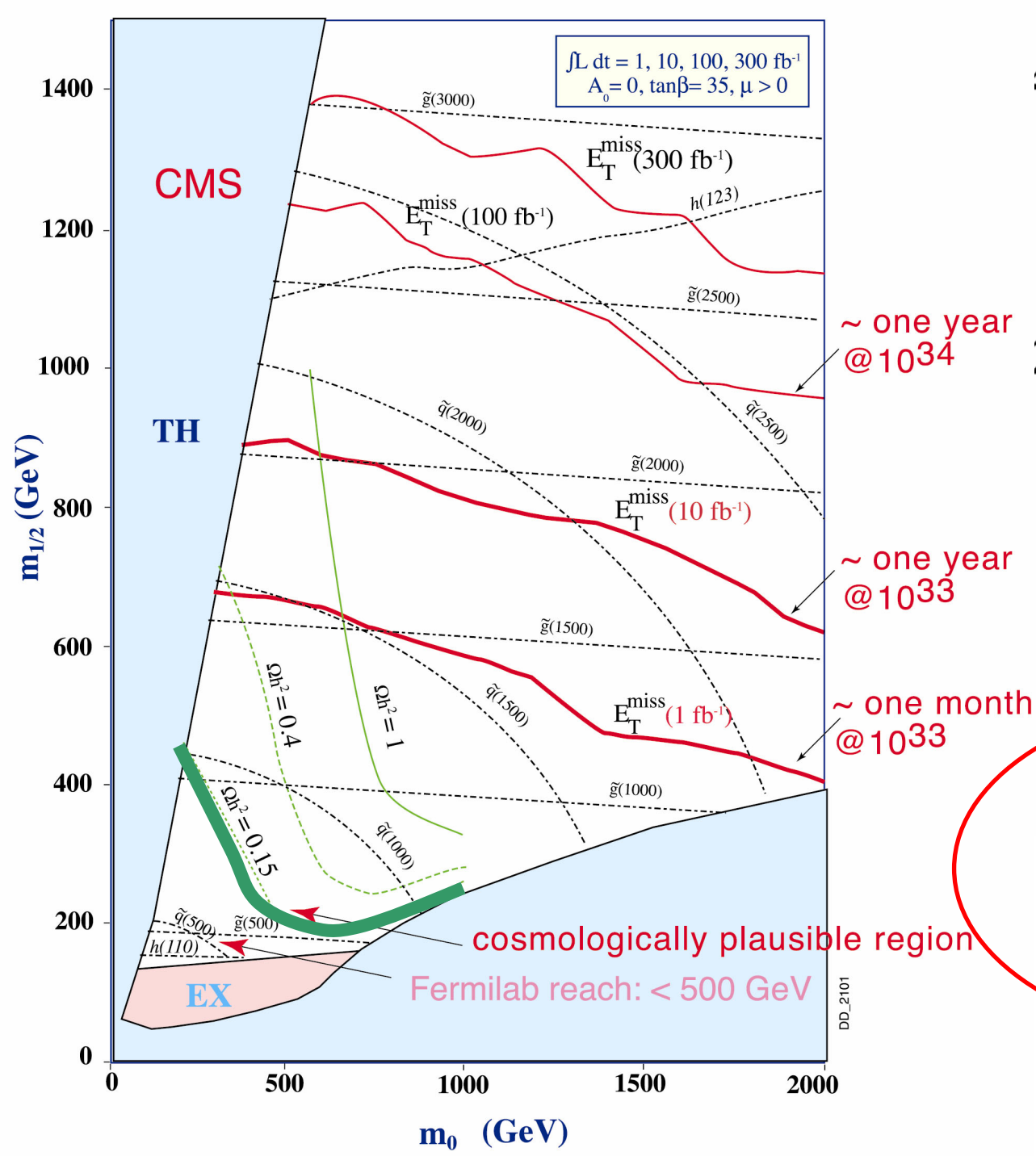


$$\Omega_{DM} = 0.23$$

Cosmological Constraints II

Evidence for relic WIMPS ?





If dark matter:
 Fermilab: **NO**
 LHC: **YES**

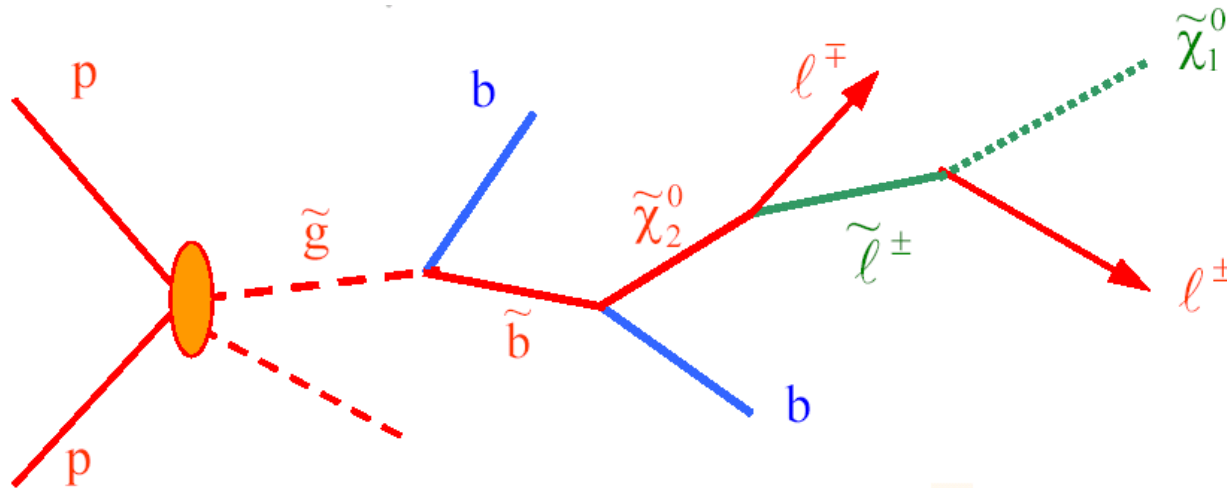
SUSY reach

SUSY search in pp

strategy:

high SM QCD background: jets

need something beyond: **leptons and/or missing energy**



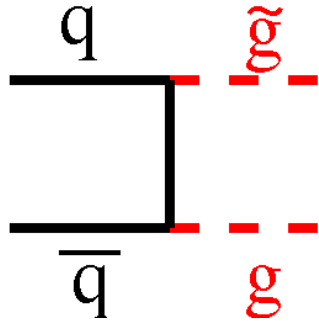
Examples:

- squarks and gluinos (missing energy) **strong**
- neutralinos and charginos (leptons and missing energy)

electroweak

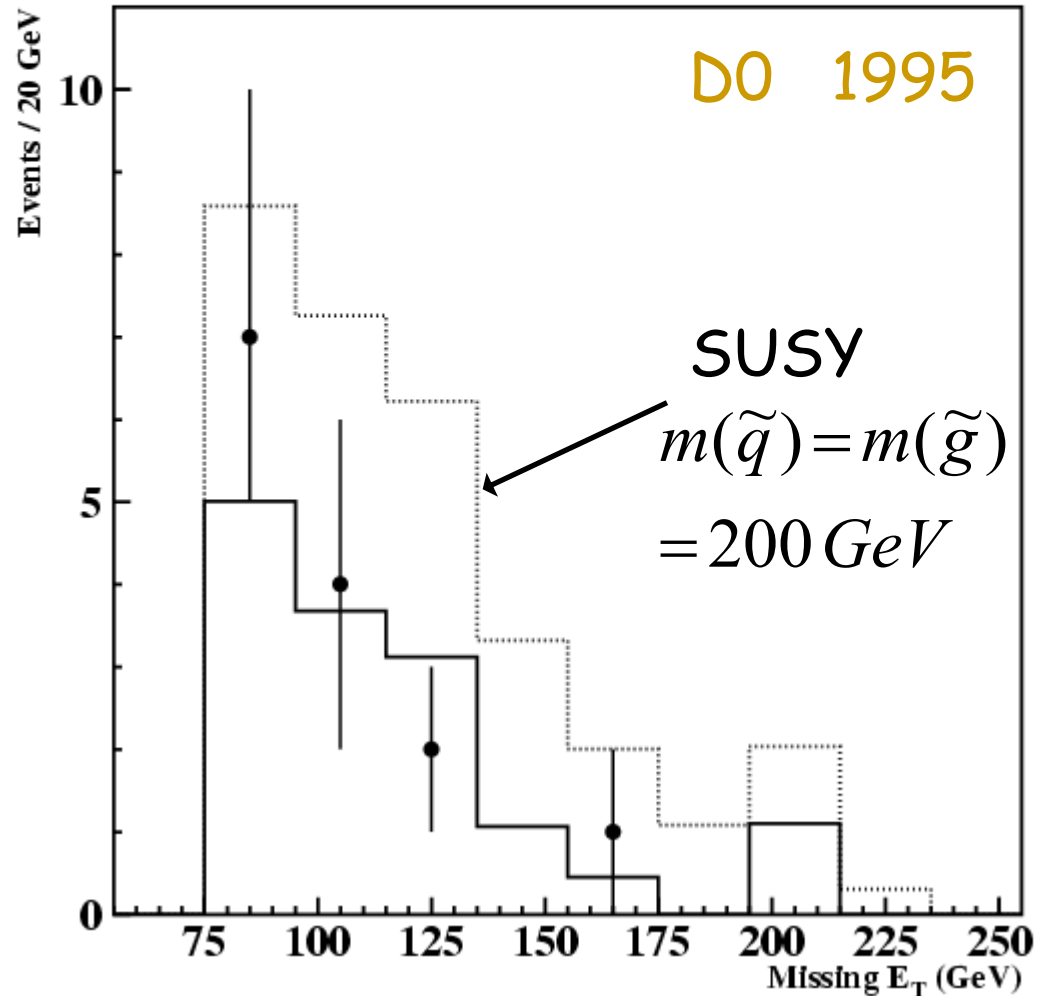
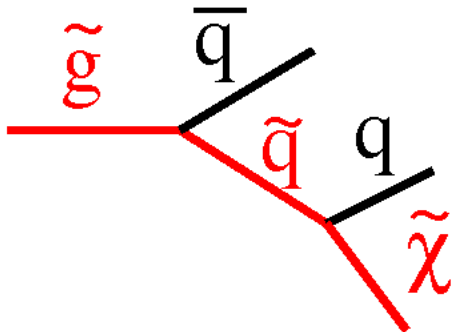
Jets + Missing Energy: Squark and Gluino Search

Production (ex.):



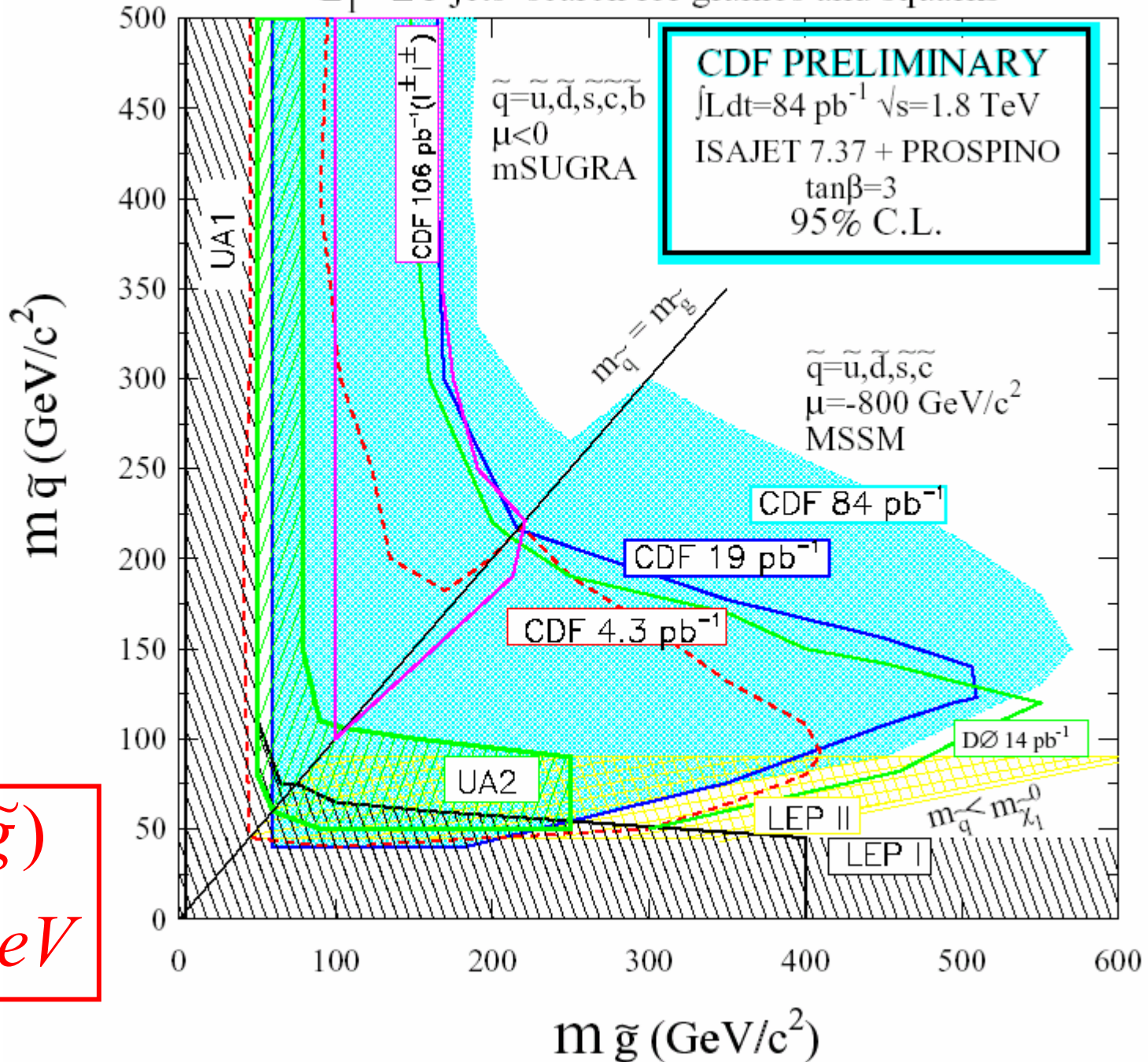
cross section large!

Decay (ex.):



Results

$E_{T+} \geq 3$ jets search for gluinos and squarks

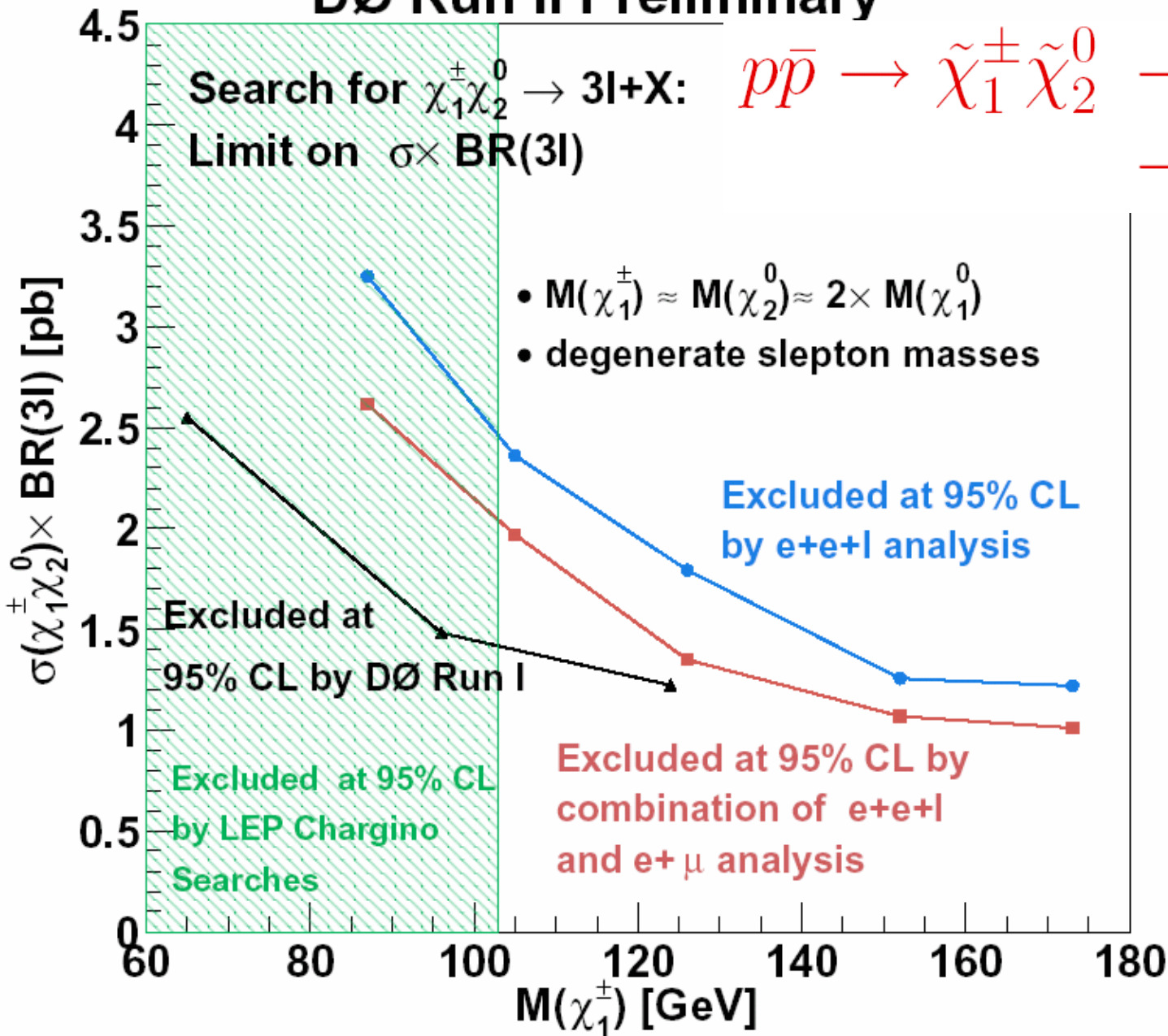
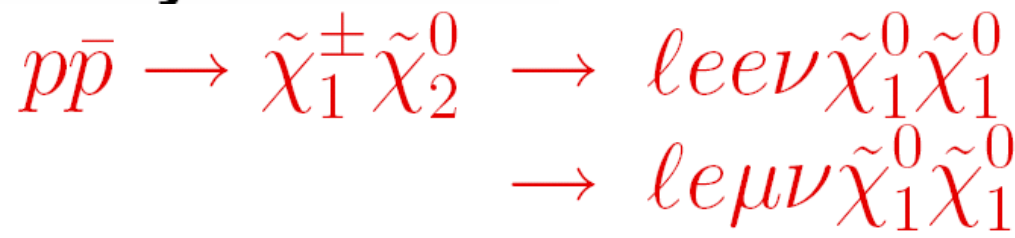


$m(\tilde{q}), m(\tilde{g})$
 $> \sim 200 \text{ GeV}$

Leptons: Hunting Charginos and Neutralinos

DØ Run II Preliminary

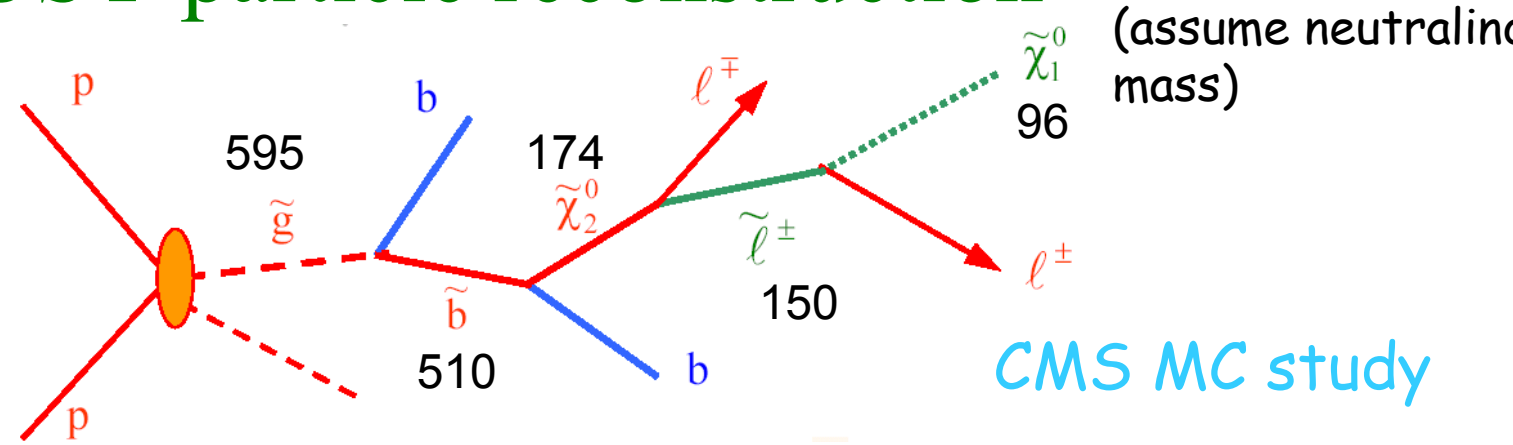
Search for $\chi_1^\pm \chi_2^0 \rightarrow 3l+X$:
 Limit on $\sigma \times \text{BR}(3l)$



Expected SUSY xsection factor ~10 lower

No (improved) SUSY limit yet

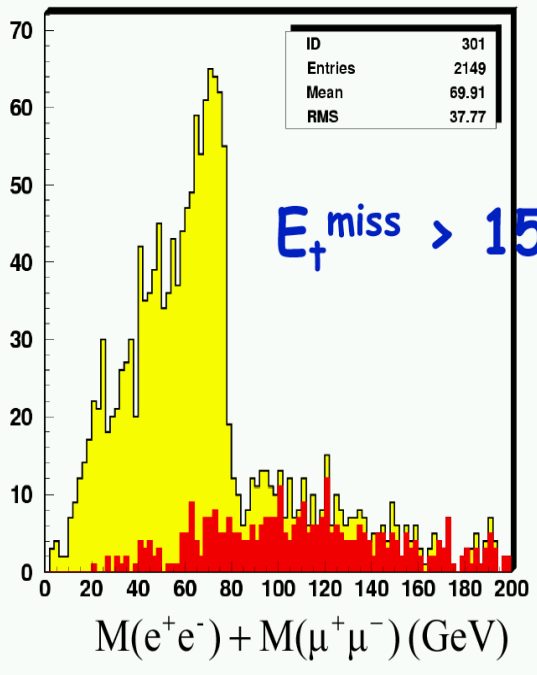
SUSY particle reconstruction



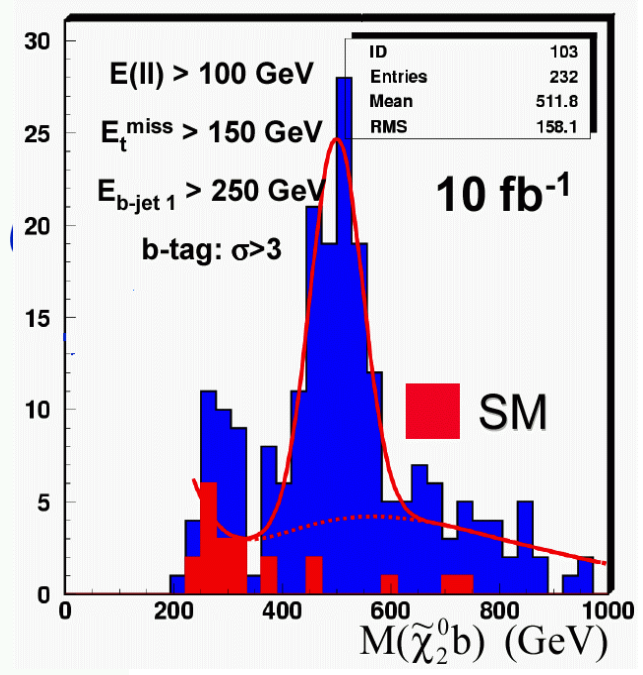
(neutralino 2)

sbottom

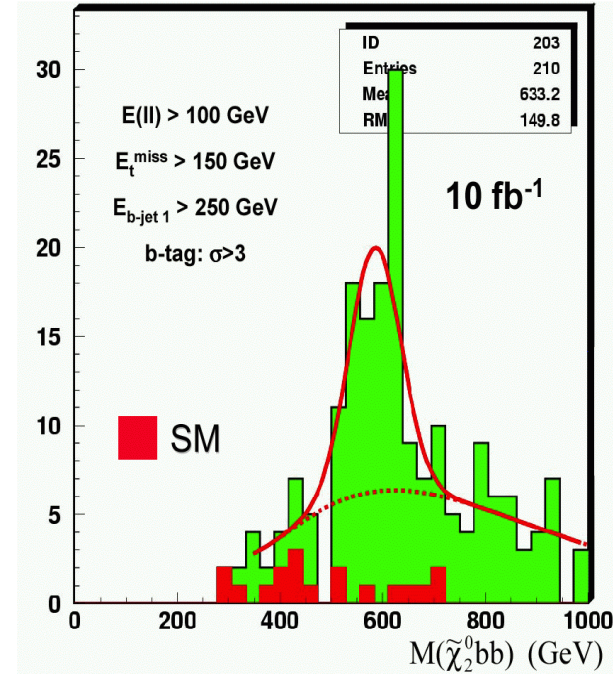
gluino



edge 78.9 +/- 2.1 GeV



499.4 +/- 6.6 GeV



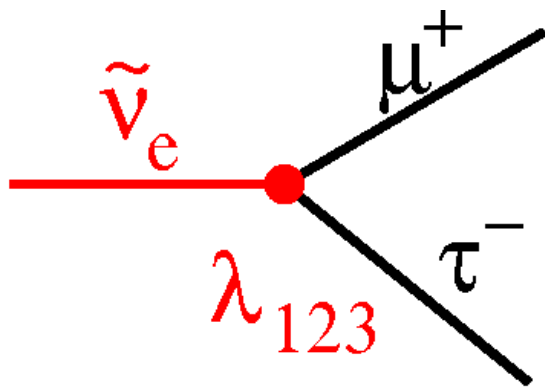
585.1 +/- 11.1 GeV

Chiorboli/Tricomi

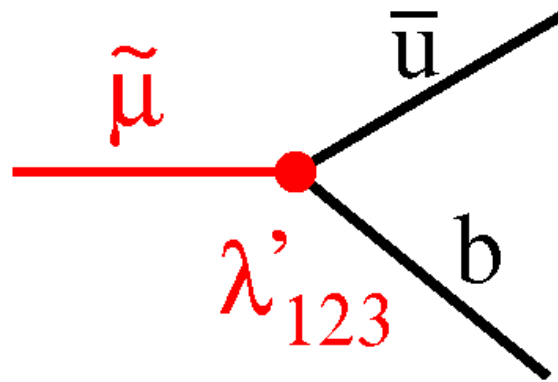
SUSY with R parity violation

- neutralino unstable, no dark matter candidate !
- lepton and/or baryon number violated

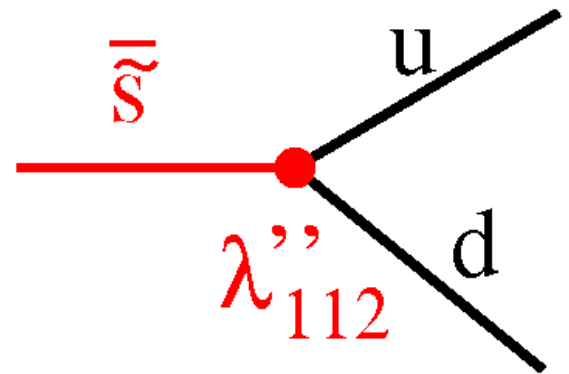
Example diagrams:



violates L



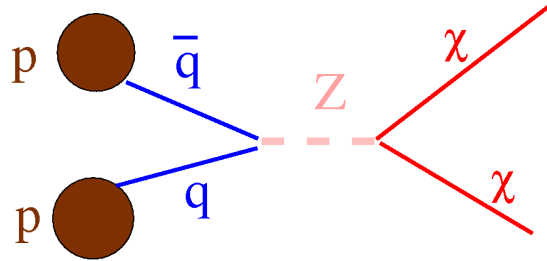
violates L



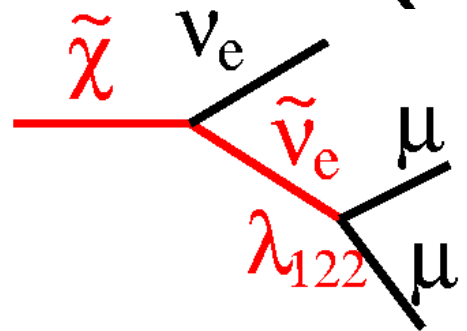
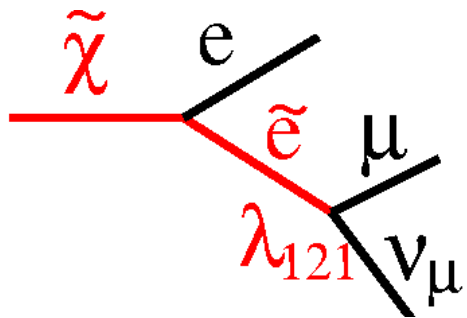
violates B

Example: Neutralino Decays via $\lambda_{121}, \lambda_{122}, \dots$

production:



decay (examples):

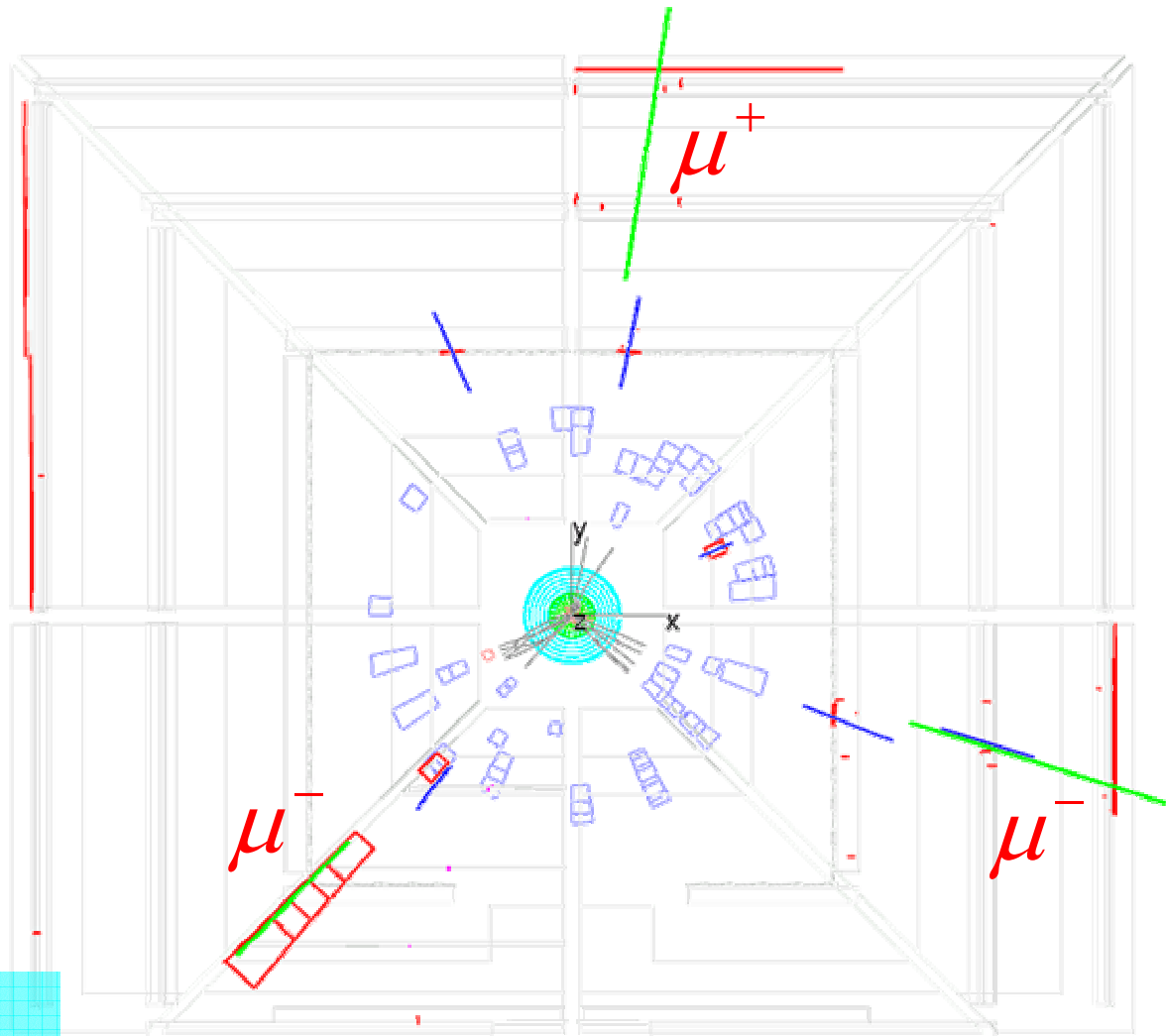


signature:

at least 3 charged leptons!

Run 170246 Event 17918459 Tue Mar 4 18:23:36 2003

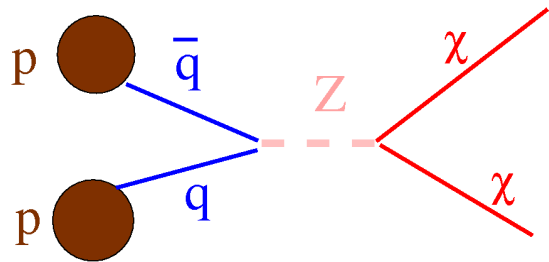
DO Run II



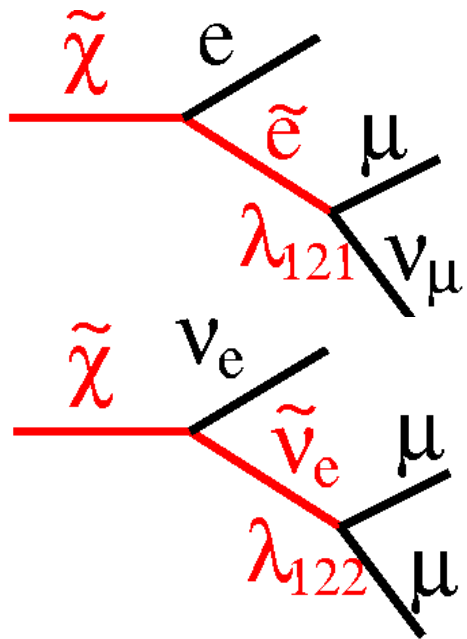
View 1, Front(X-Y)

Example: Neutralino Decays via $\lambda_{121}, \lambda_{122}, \dots$

production:



decay (examples):

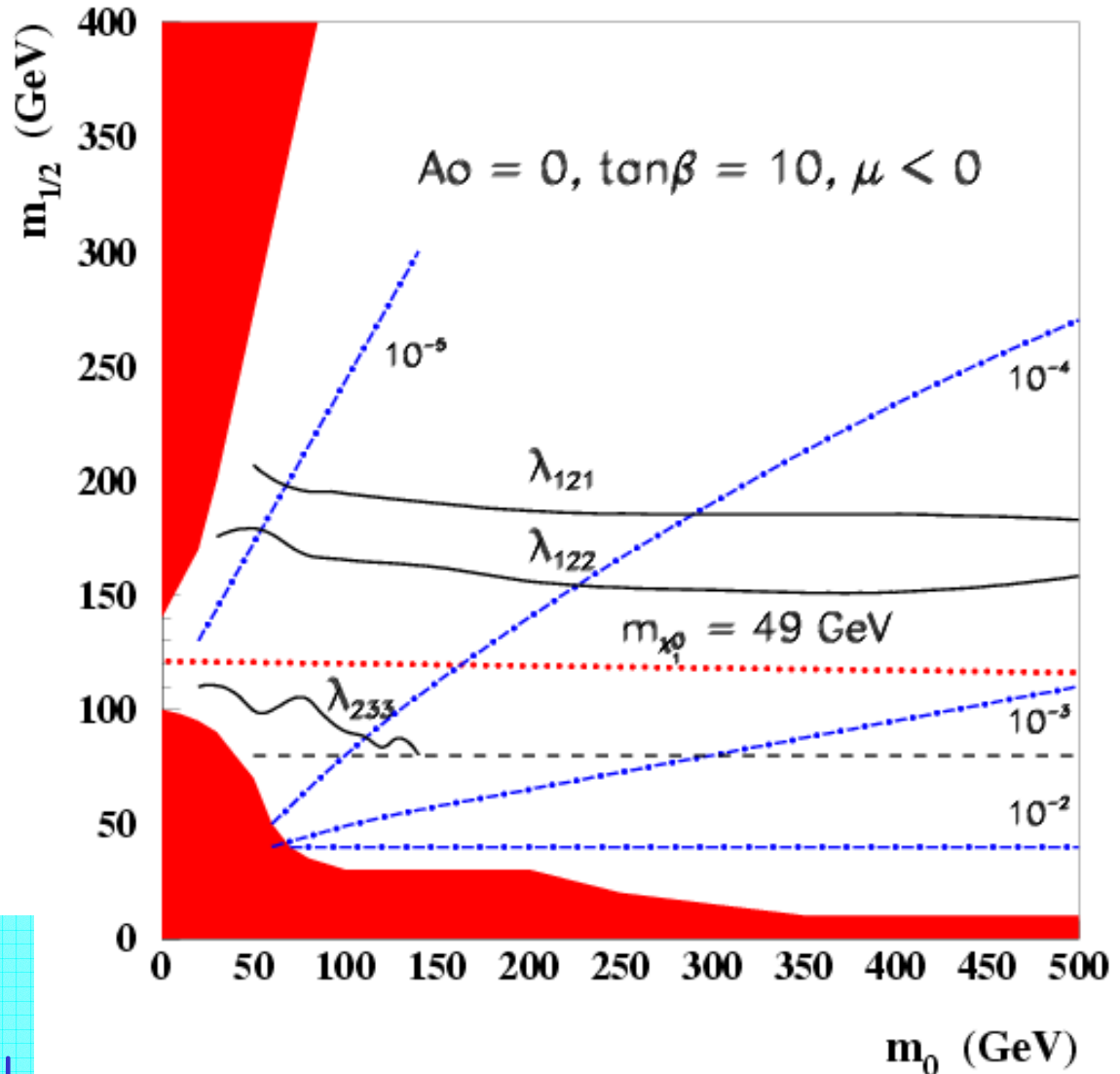


experimental signature:

at least 3 charged leptons!

upper limits:

D0 Run



(Large) Extra Dimensions

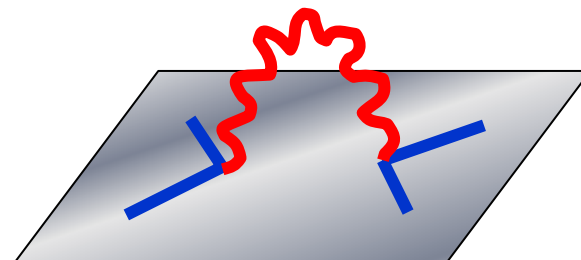
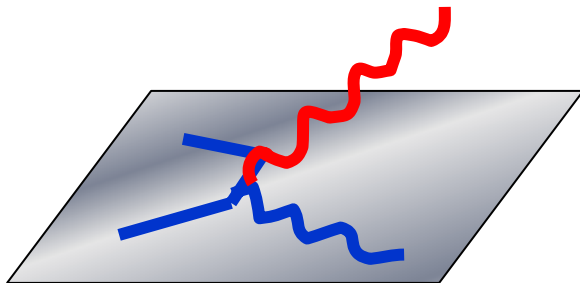
Why is **gravity** so different from the other interactions ?

mass and length scales:

$$M_{ew} \sim 10^2 \text{ GeV} \quad l_{ew} \sim 10^{-18} \text{ m}$$

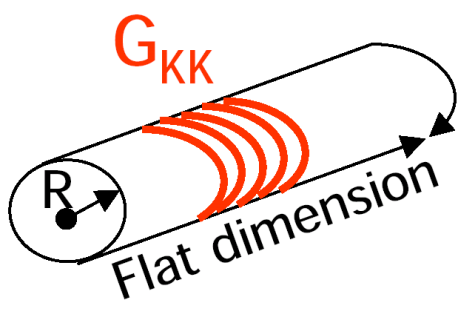
$$M_{Pl} \sim \frac{1}{\sqrt{G_N}} \sim 10^{19} \text{ GeV} \quad l_g \sim 10^{-35} \text{ m}$$

Idea: only one fundamental scale $M_S \sim 100 - 1000 \text{ GeV}$
 gravity appears weak since gravitons propagate
 in $4 + n$ dimensions („dilution“)



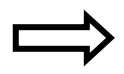
Extra Dimensions: phenomenology

n extra dimensions of space with size R:



Compactified dimension

$$V \sim \frac{1}{M_{Pl}^2} \frac{m_1 m_2}{r}$$



$$V \sim \frac{1}{M_S^{2+n}} \frac{1}{R^n} \frac{m_1 m_2}{r}$$



- deviations from Newton/Einstein laws for $r < R$

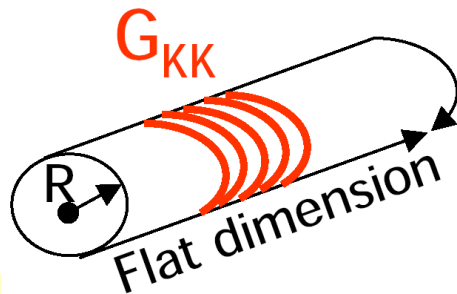
therefore n=1 and n=2 ruled out!

- gravitons G show up in high energy physics experiments as real or virtual particles

pp colliders!

Extra Dimensions: phenomenology

n extra dimensions of space with size R:



Compactified
dimension

$$R = \frac{1}{2\sqrt{\pi}M_S} \left(\frac{M_{Pl}}{M_S} \right)^{2/n}$$

$$8 \times 10^{12} \text{ m}, \quad n = 1$$

$$0.7 \text{ mm}, \quad n = 2$$

$$3 \text{ nm}, \quad n = 3$$

$$6 \times 10^{-12} \text{ m}, \quad n = 4$$



Escher

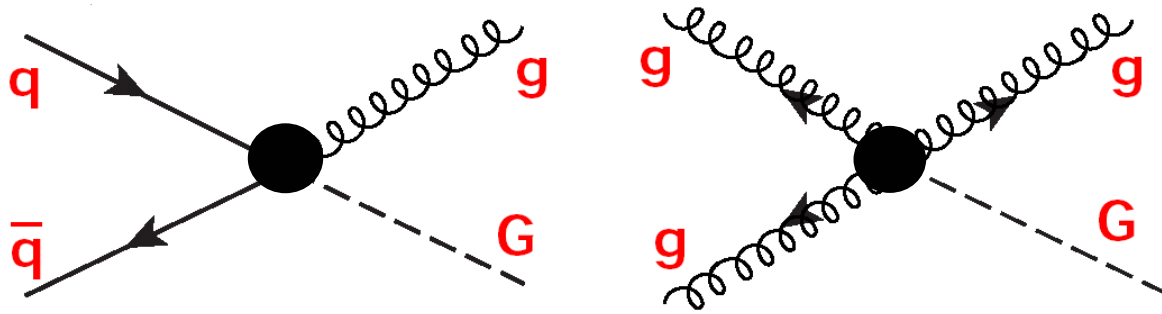
- deviations from Newton/Einstein laws for $r < R$

therefore n=1 and n=2 ruled out!

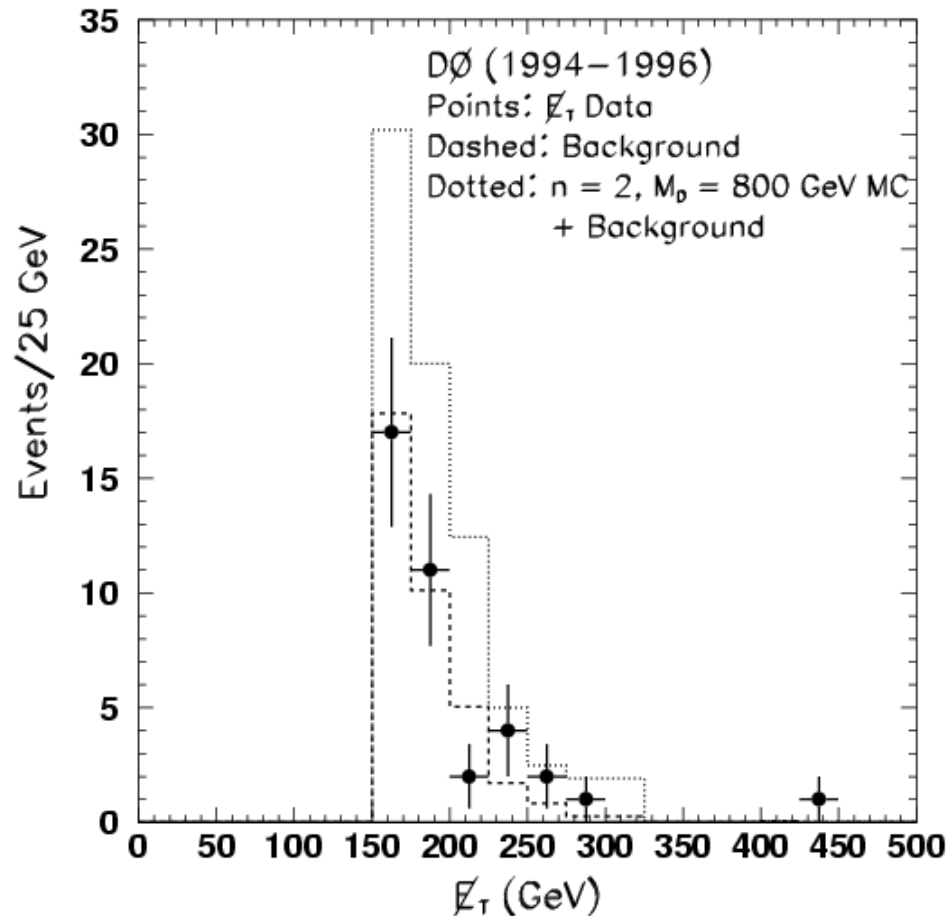
- gravitons G show up in high energy physics experiments as real or virtual particles

pp colliders!

Real Graviton Emission in $p p$



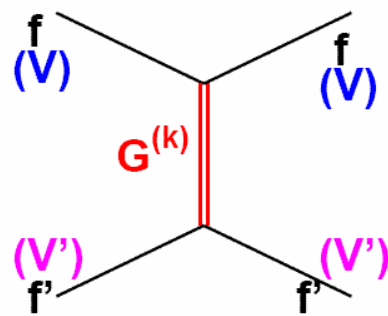
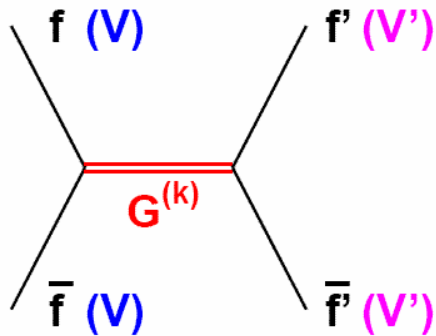
Monojet signatures!



resulting lower limits on M_S/GeV :

$n=2$	$n=3$	$n=4$	$n=5$	$n=7$
890	730	680	640	620

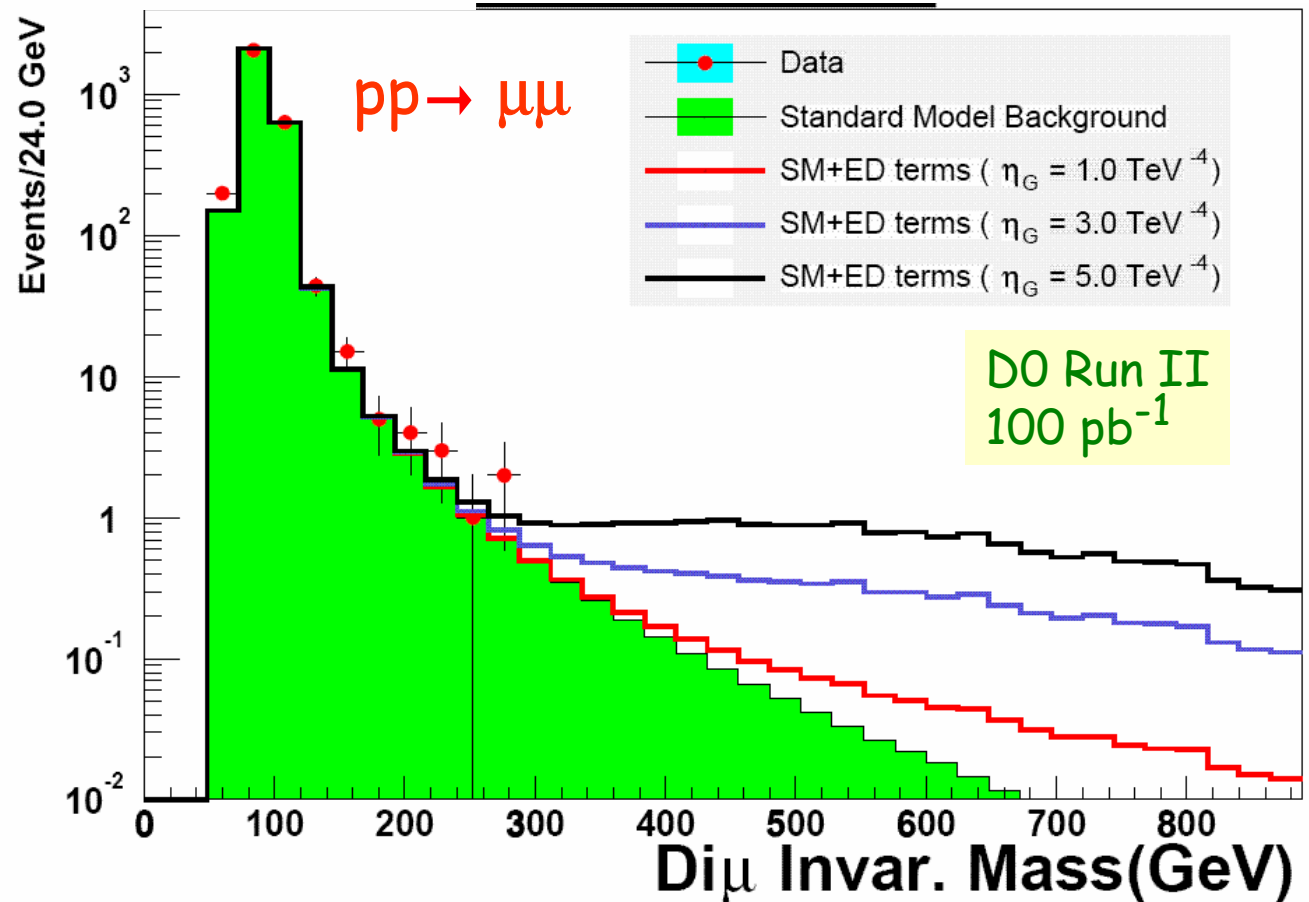
Virtual Graviton effects in $p p$



SM cross sections modified!

All D0 analyses
($ee, \mu\mu, \gamma\gamma$)
combined:

$$M_S (n=2) > 1.38 \text{ TeV}$$



Black holes ?

predicted in large extra dimension models

production: mass 1 - 10 TeV, xsection large (\sim nb)

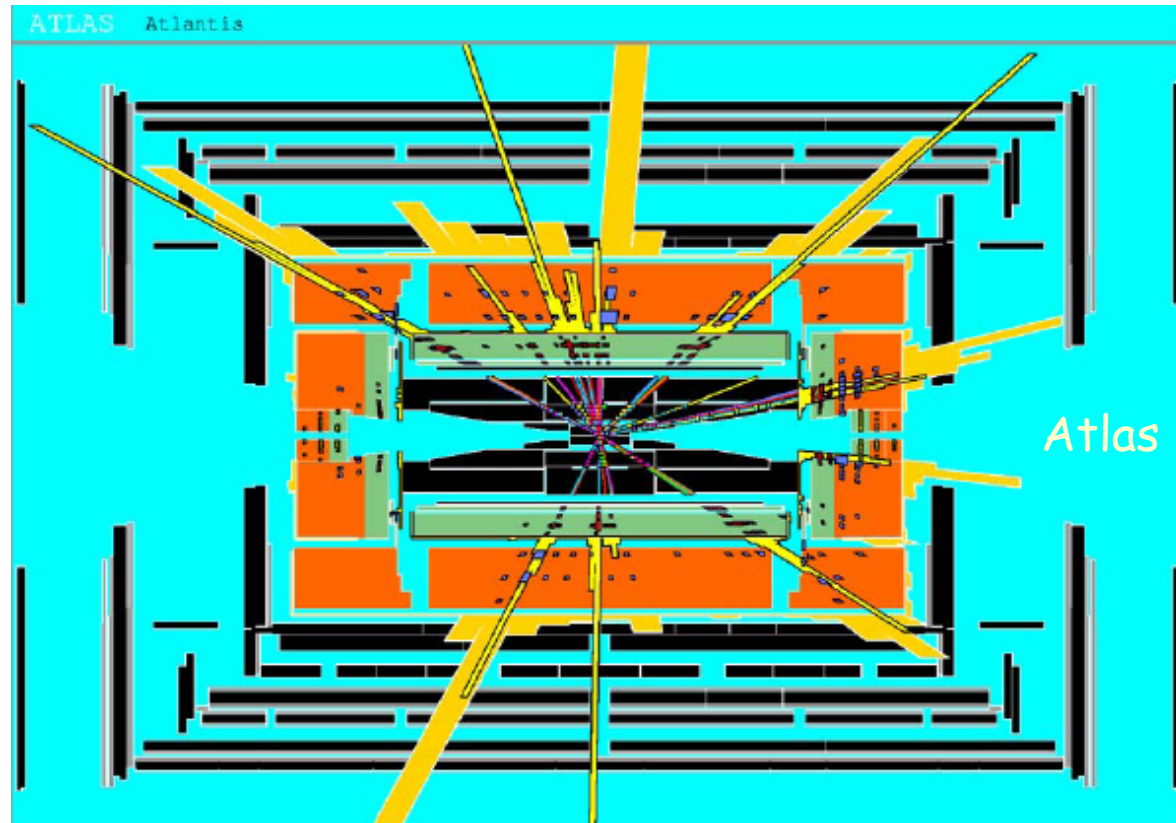
decay:

Hawking radiation

All SM d.o.f. equally likely

Multiplicity up to 30

Decay also into higgs!



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 - searches:
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 - R-Parity violated
- Extra dimensions
- Black holes

References

References

- lectures:

F. Gianotti, LHC physics

www.wlap.org/cern/lectures/summer/2000/gianotti

J. Womersley, Physics at Hadron Colliders

d0server1.fnal.gov/users/womersley/brazil1.pdf...brazil4.pdf

- experimental homepages:

www-cdf.fnal.gov

www-d0.fnal.gov

atlas.web.cern.ch

cmsinfo.cern.ch

- theory:

Physics at Run II workshop

fnth37.fnal.gov/run2.html