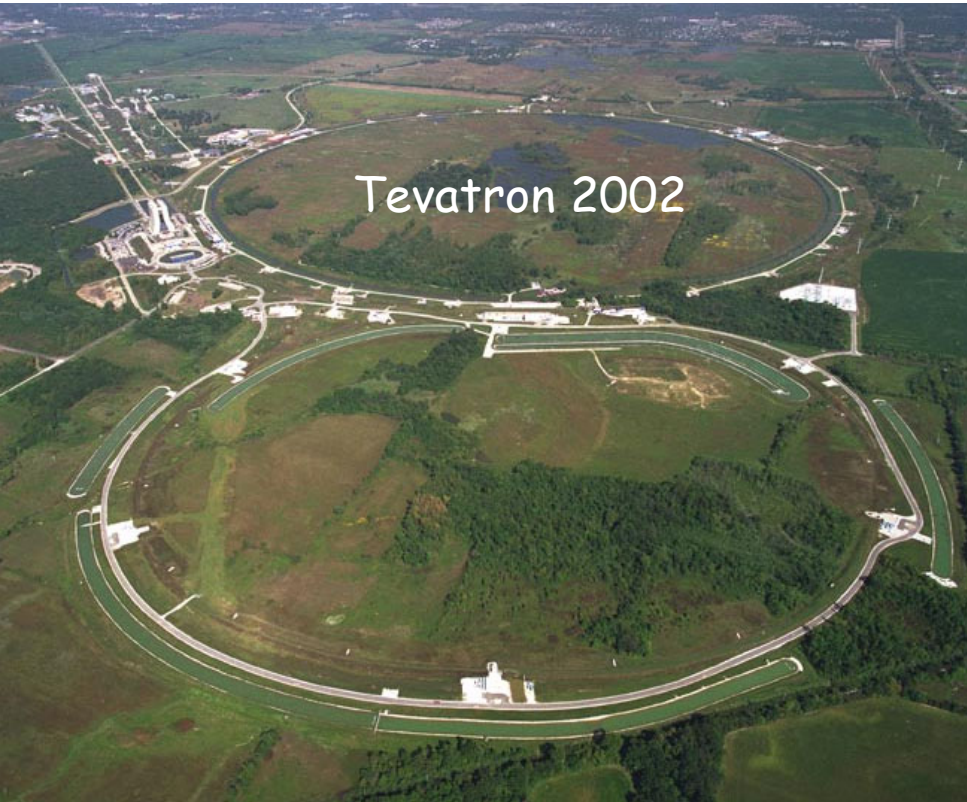
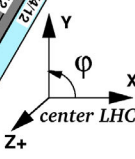
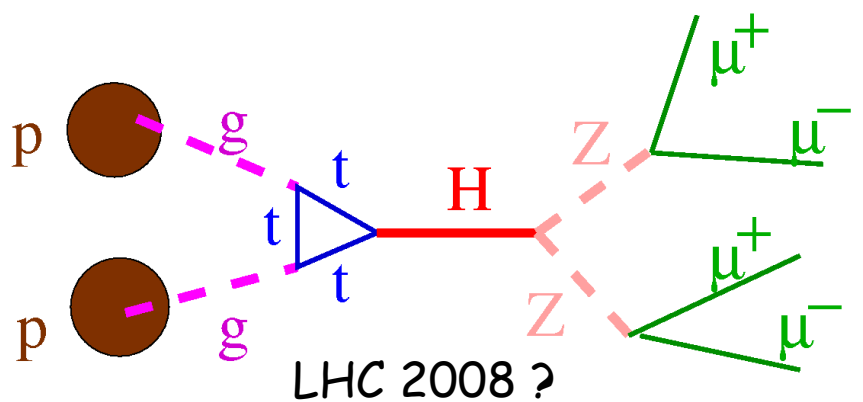
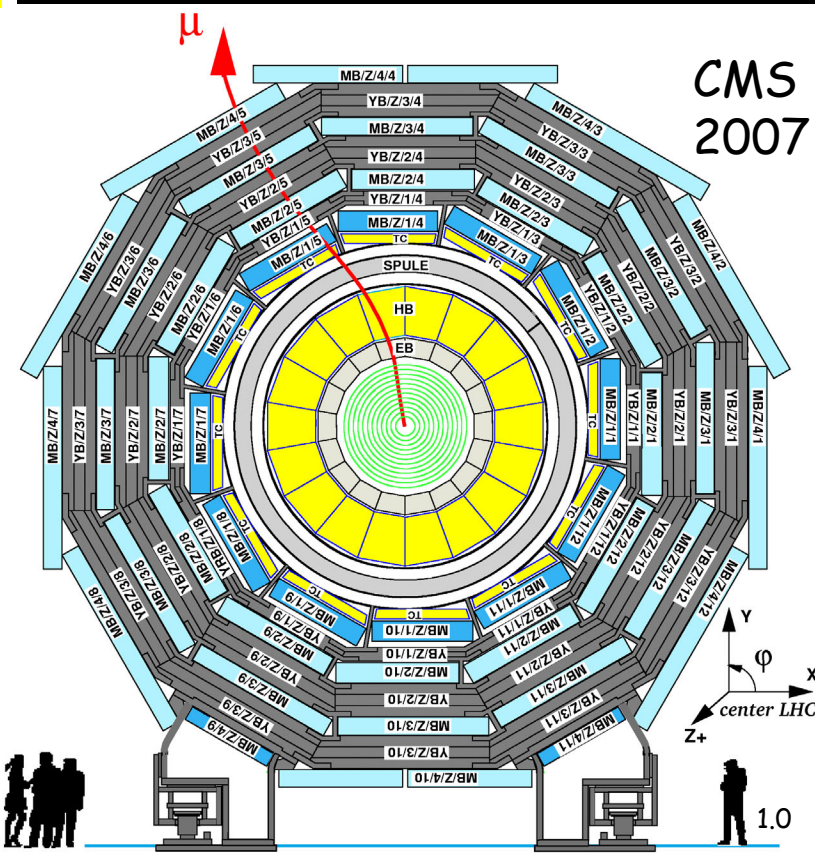
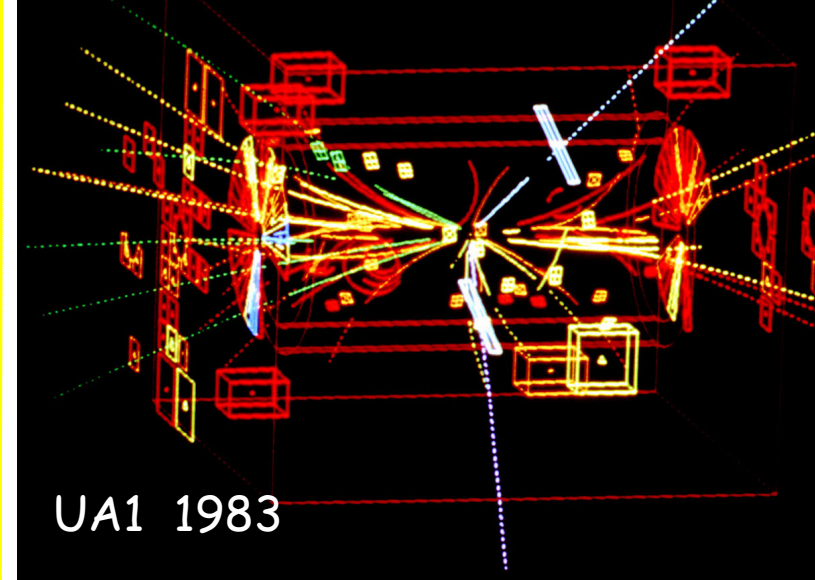


part IV



p
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y
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s



1.0

Part I Introduction

Part II Standard Model Physics

Part III Higgs

Part IV New Phenomena

- SUSY
 - motivation
 - searches:
 - R-Parity conserved
 - R-Parity violated
- Extra dimensions
- Black holes

References

SUperSymmetry

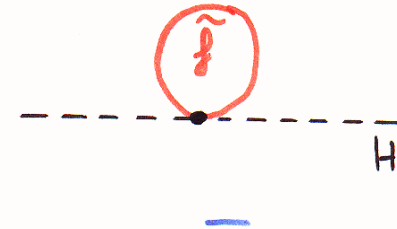
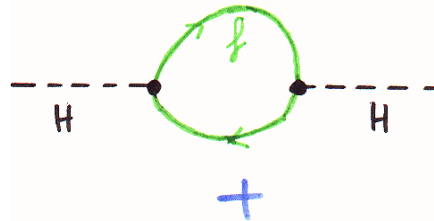
Particle	Spin	Susy-Partner	Spin
ν_e	1/2	$\tilde{\nu}_e^L$	0
e^-	1/2	$\tilde{e}_L^-, \tilde{e}_R^-$	0
u	1/2	\tilde{u}_L, \tilde{u}_R	0
d	1/2	\tilde{d}_L, \tilde{d}_R	0
γ, Z, h, H, A	1, 0	$\tilde{\chi}_1^0, \tilde{\chi}_2^0, \tilde{\chi}_3^0, \tilde{\chi}_4^0$	1/2
W^\pm, H^\pm	1, 0	$\tilde{\chi}_1^\pm, \tilde{\chi}_2^\pm$	1/2
g	1	\tilde{g}	1/2

if R-parity (-1 for sparticles) is conserved:
 \Rightarrow LSP = Lightest SUSY particle = $\tilde{\chi}_1$ = stable

SUSY

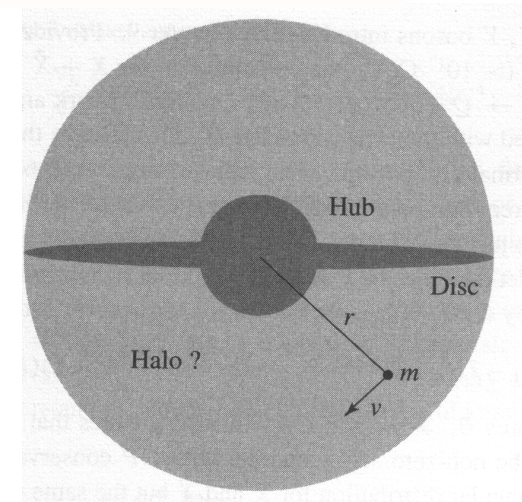
Nice features:

- symmetry relating **bosons (forces)** \leftrightarrow fermions (matter)
- higgs mass m_H
under control
- grand unification (incl. gravity!) possible
- neutralino = dark matter candidate



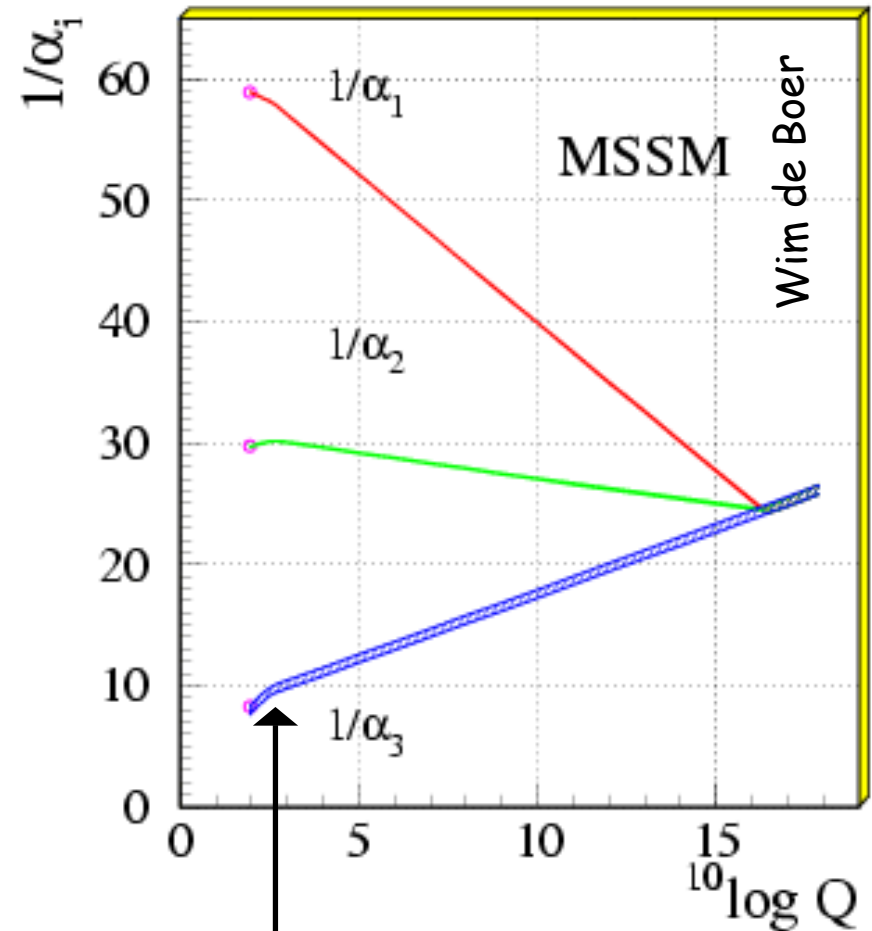
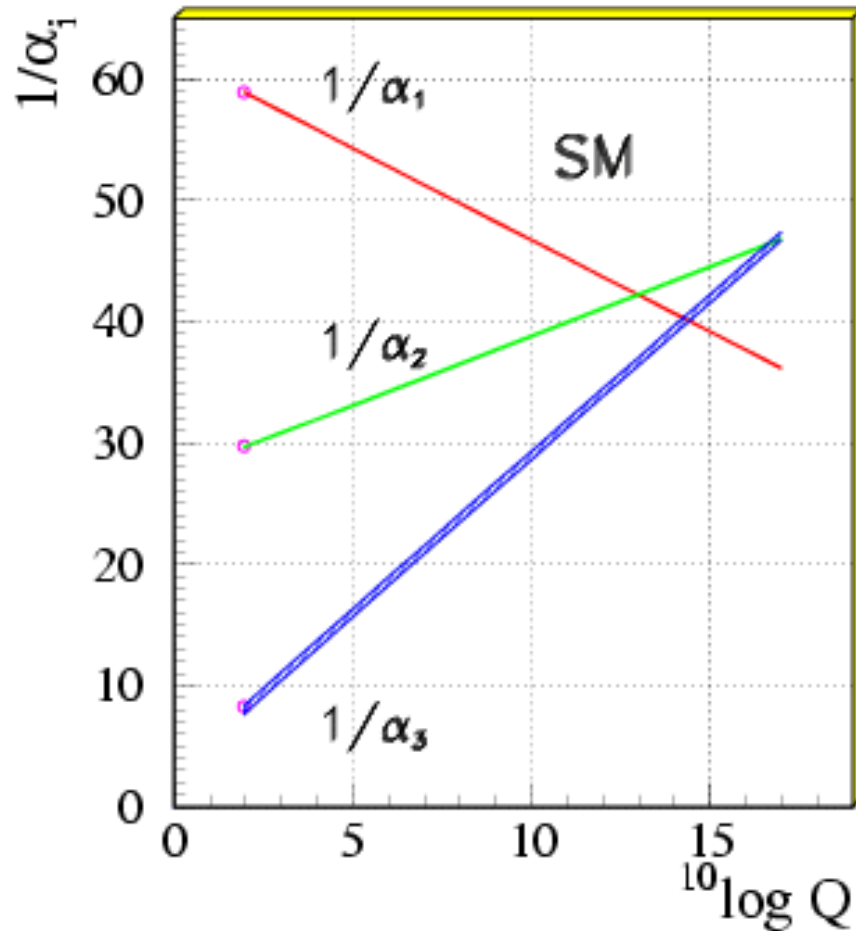
Other properties:

- SUSY broken: **no sparticle seen yet**
- > 100 new parameters \Rightarrow **Minimal Model (MSSM, MSUGRA)**



Grand Unification ?

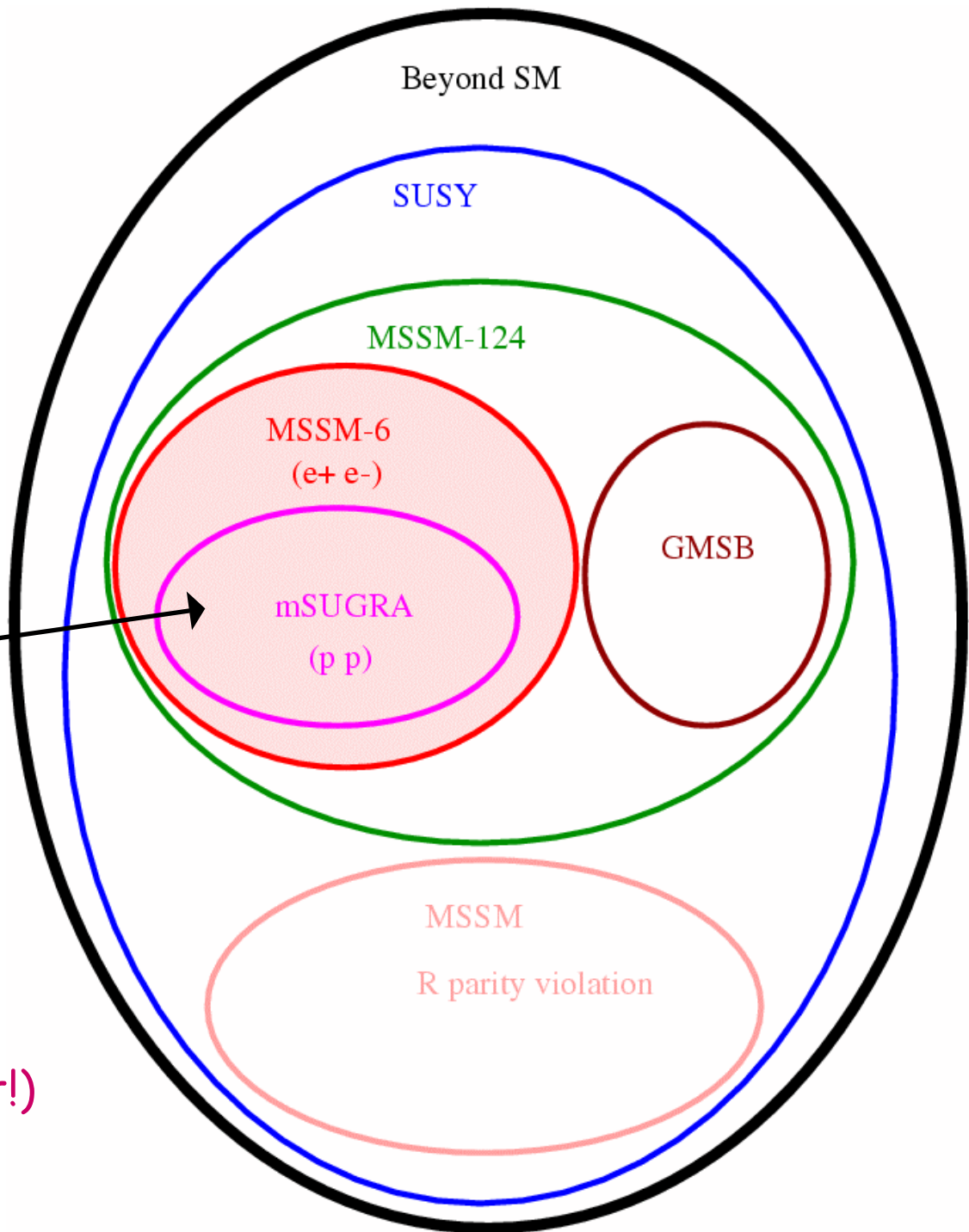
Unification of the Coupling Constants in the SM and the minimal MSSM



SUSY mass scale ~ TeV

Wim de Boer

SUSY models



investigated
most often
by pp experiments

(does not mean it's right!)

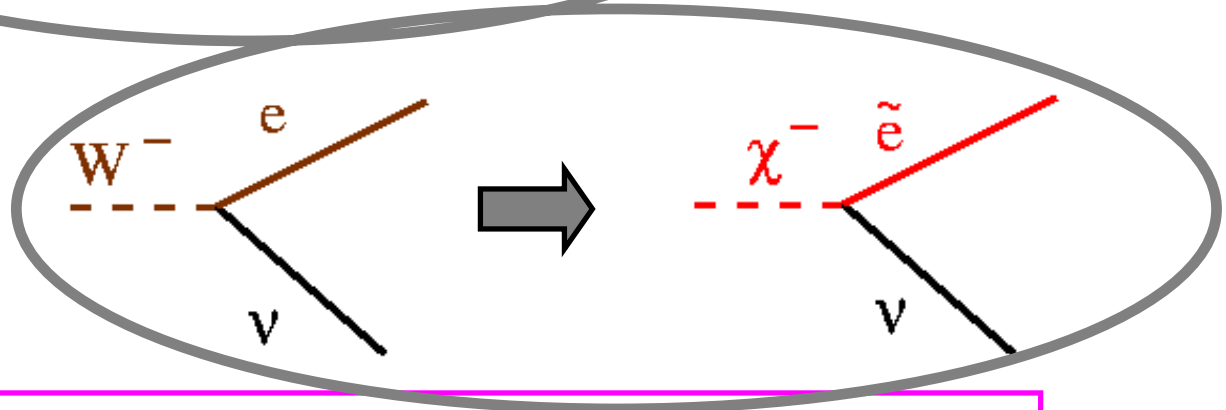
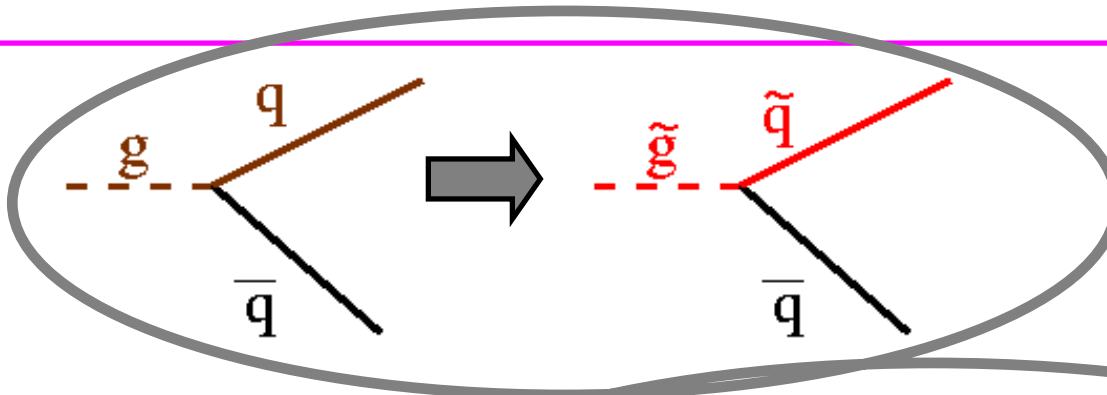
SUSY interactions (R conserved)

Feynman graphs:

take any SM vertex with 3 or 4 particles

replace two legs by the corresponding sparticles (\sim)

Examples



Coupling constants (electroweak, strong):

same as in SM !

MSUGRA parameters

MSUGRA = Minimal SuperGRAvity model

- m_0 = universal scalar mass at GUT scale (s..., higgs)
- $m_{1/2}$ = universal gaugino mass at GUT scale (...inos)
- $\tan \beta$ = v_2/v_1 = ratio of higgs vacuum expectation values
- A_0 = universal sfermion mass mixing parameter [GUT]
- $\text{sgn } \mu$ = sign of higgsino mass parameter

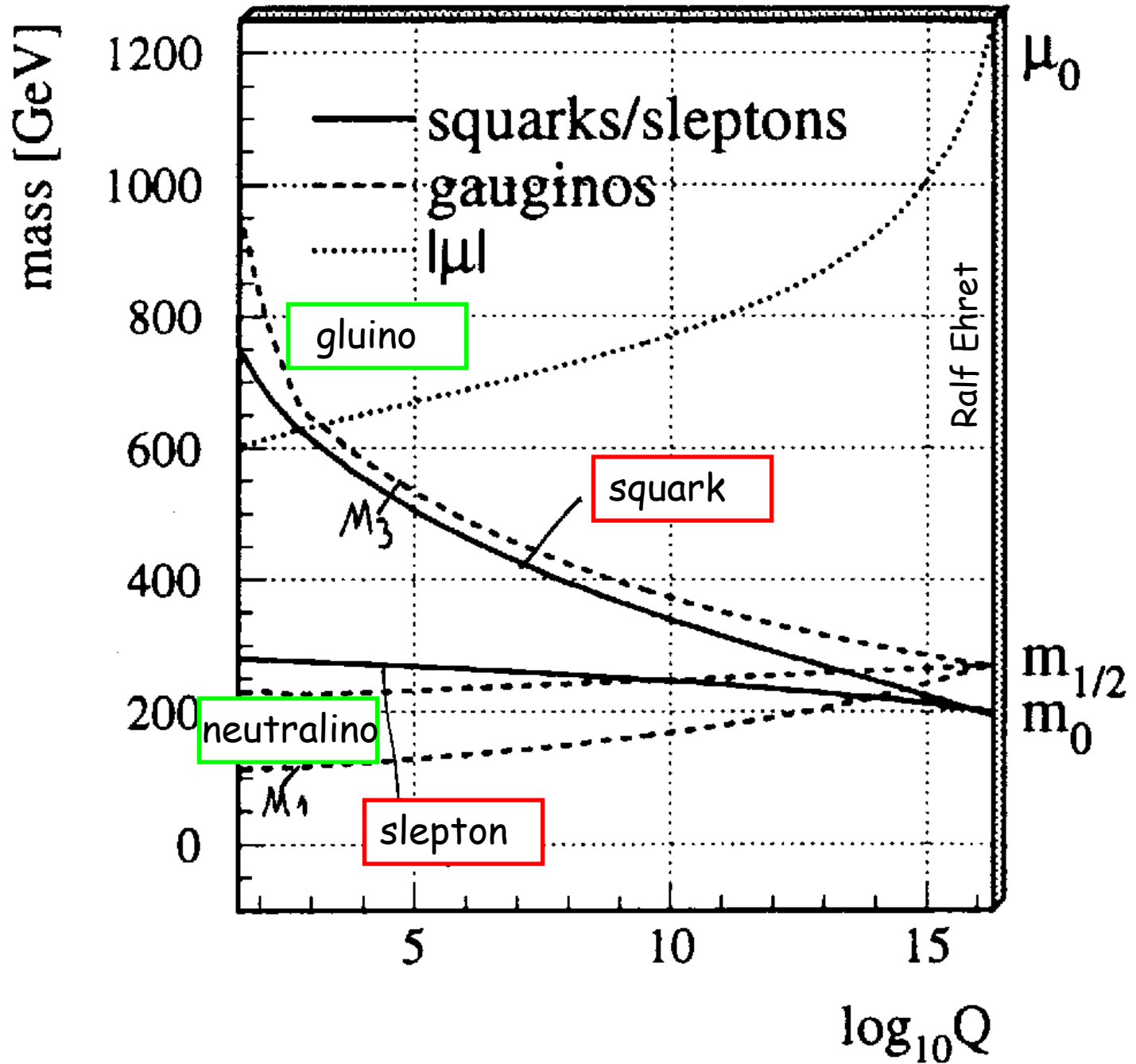
Required:

$M(\text{SUSY}) < 1 \text{ TeV}$

LSP without electromagnetic and strong coupling

Note: m_h given by m_0 ... LEP higgs limit = severe constraint

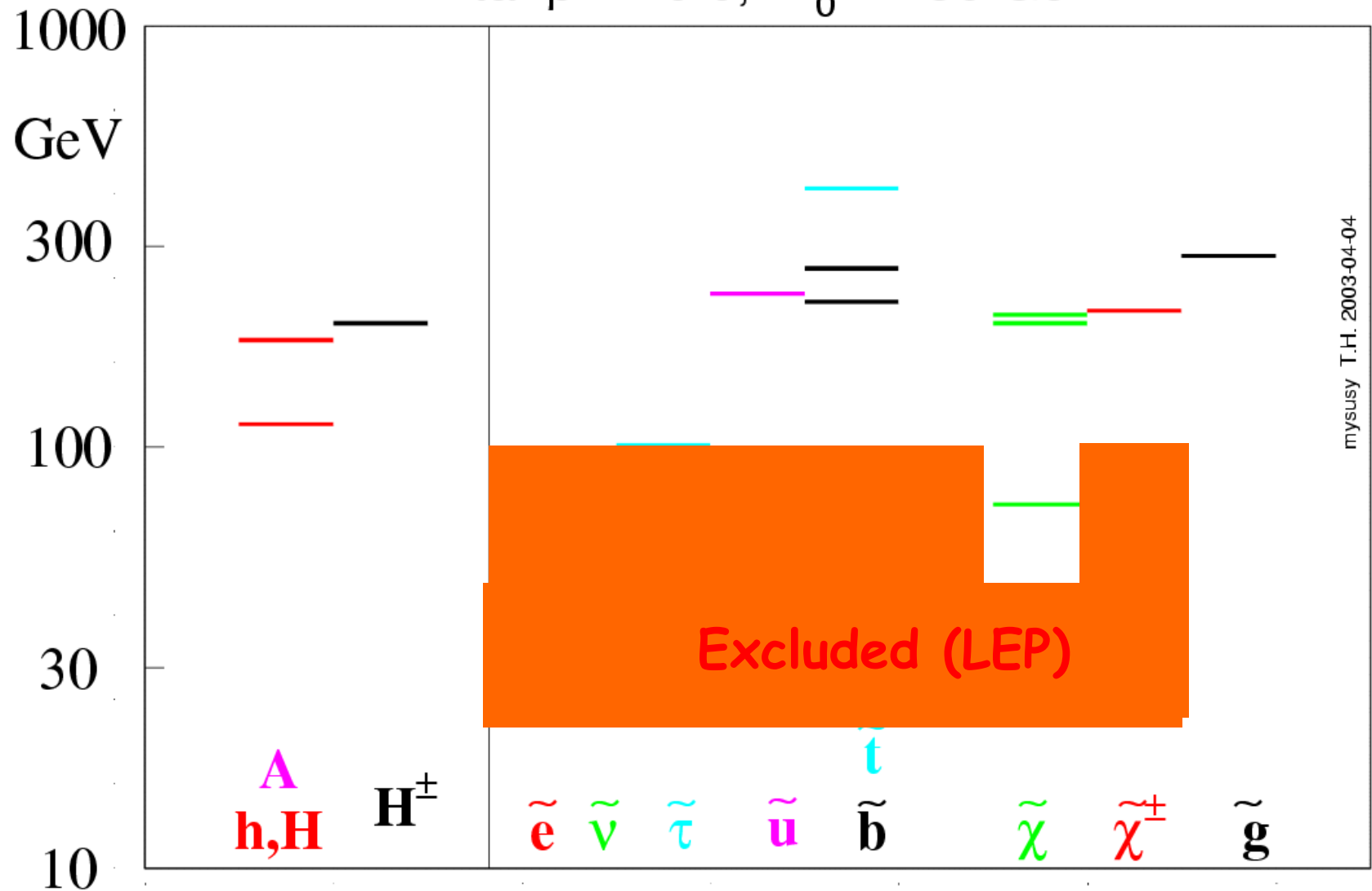
MSUGRA masses



MSUGRA scenario 1

$$m_0 = 10 \text{ GeV}, \quad m_{1/2} = 100 \text{ GeV}, \quad \mu \text{ neg.}$$

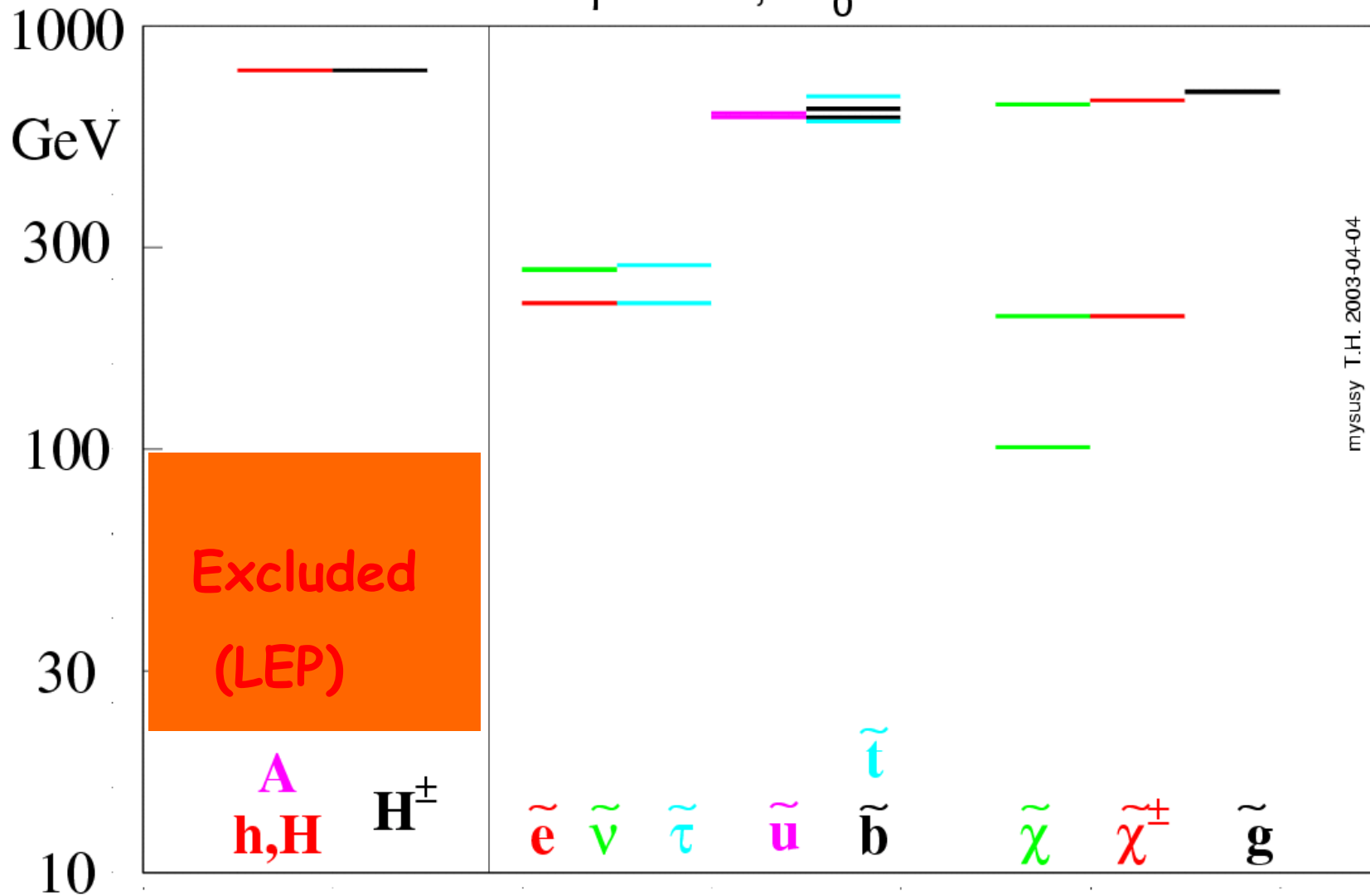
$$\tan\beta = 10.0, \quad A_0 = 450 \text{ GeV}$$

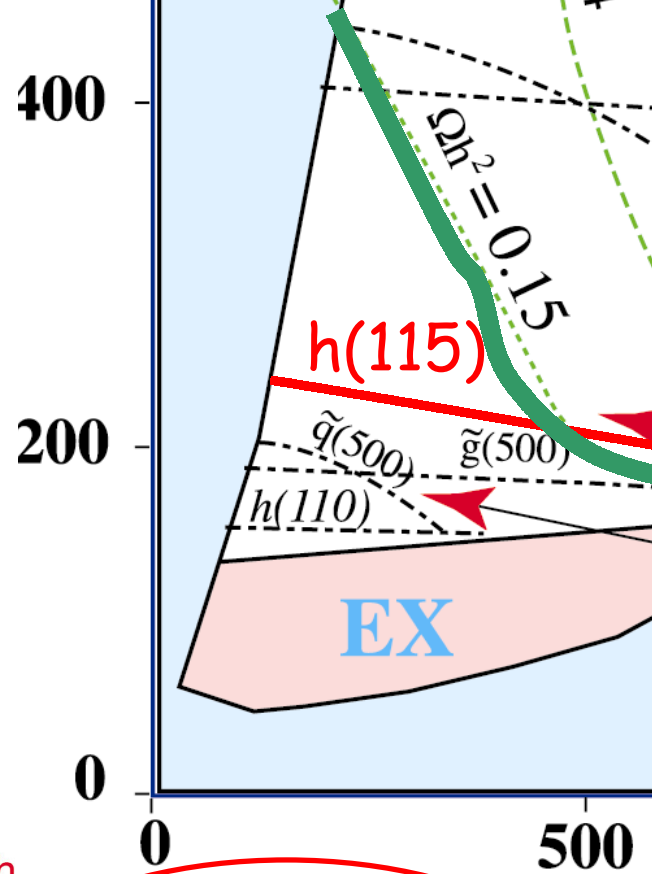
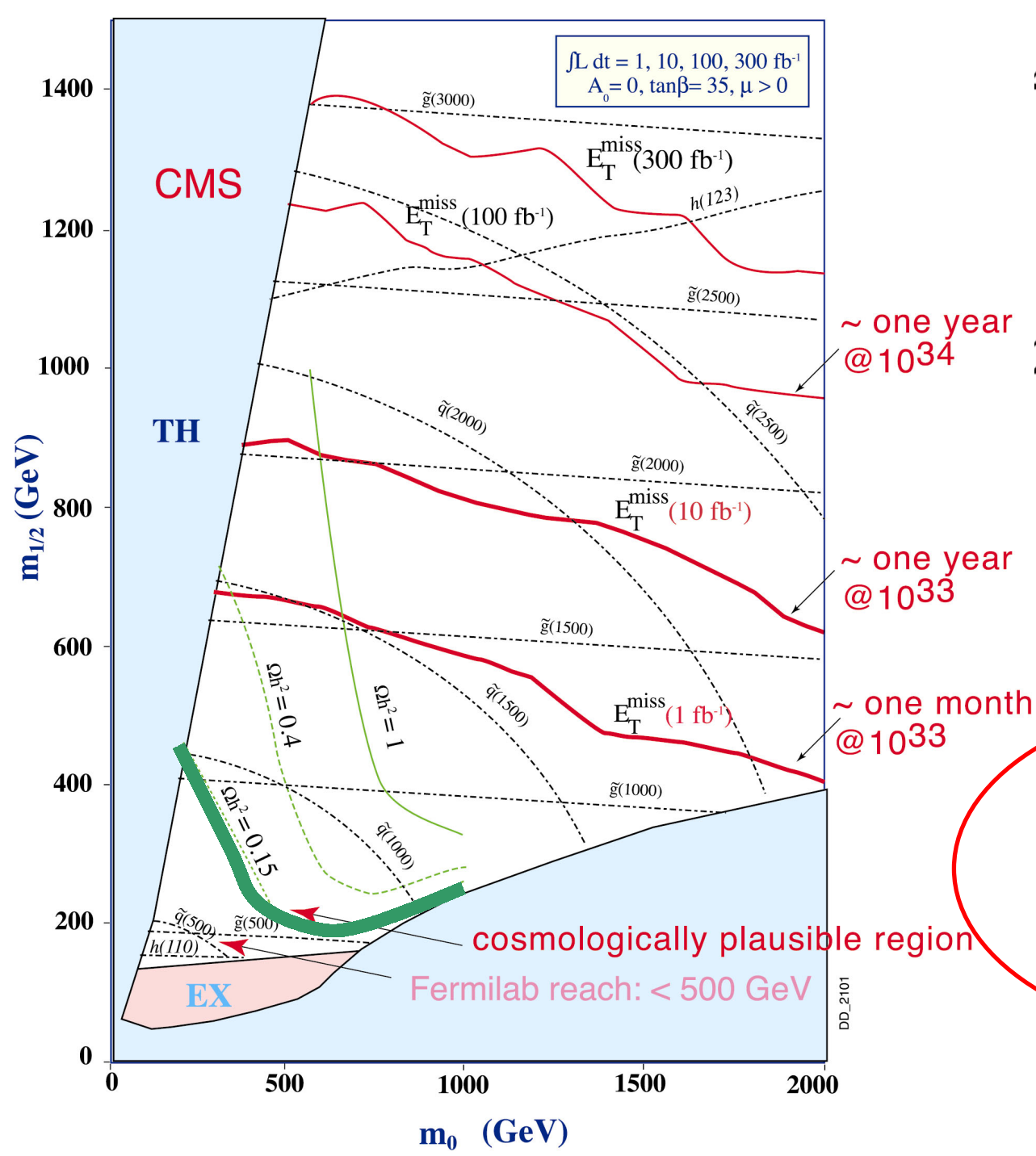


MSUGRA scenario 2

$$m_0 = 200 \text{ GeV}, \quad m_{1/2} = 243 \text{ GeV}, \quad \mu \text{ neg.}$$

$$\tan\beta = 2.0, \quad A_0 = 0 \text{ GeV}$$





Cosmological
 constraint (dark
 matter)
 essential!

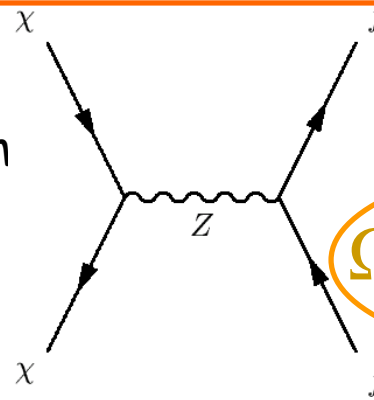
SUSY reach

Cosmological Constraints I

Assume: neutralino = $\tilde{\chi}_0$ = dark matter = WIMP

Early universe:

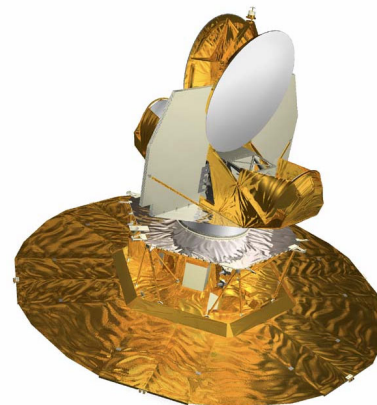
- theory
- 1) Very high temperature and pressure:
creation and annihilation: equilibrium
 - 2) High temperature and pressure:
annihilation dominates: $N(\tilde{\chi}_0) \downarrow$
 - 3) Low temperature and pressure:
freeze out: $N(\tilde{\chi}_0) = \text{const}$



$$\Omega_{DM} = f(m_\chi, \dots)$$

observation

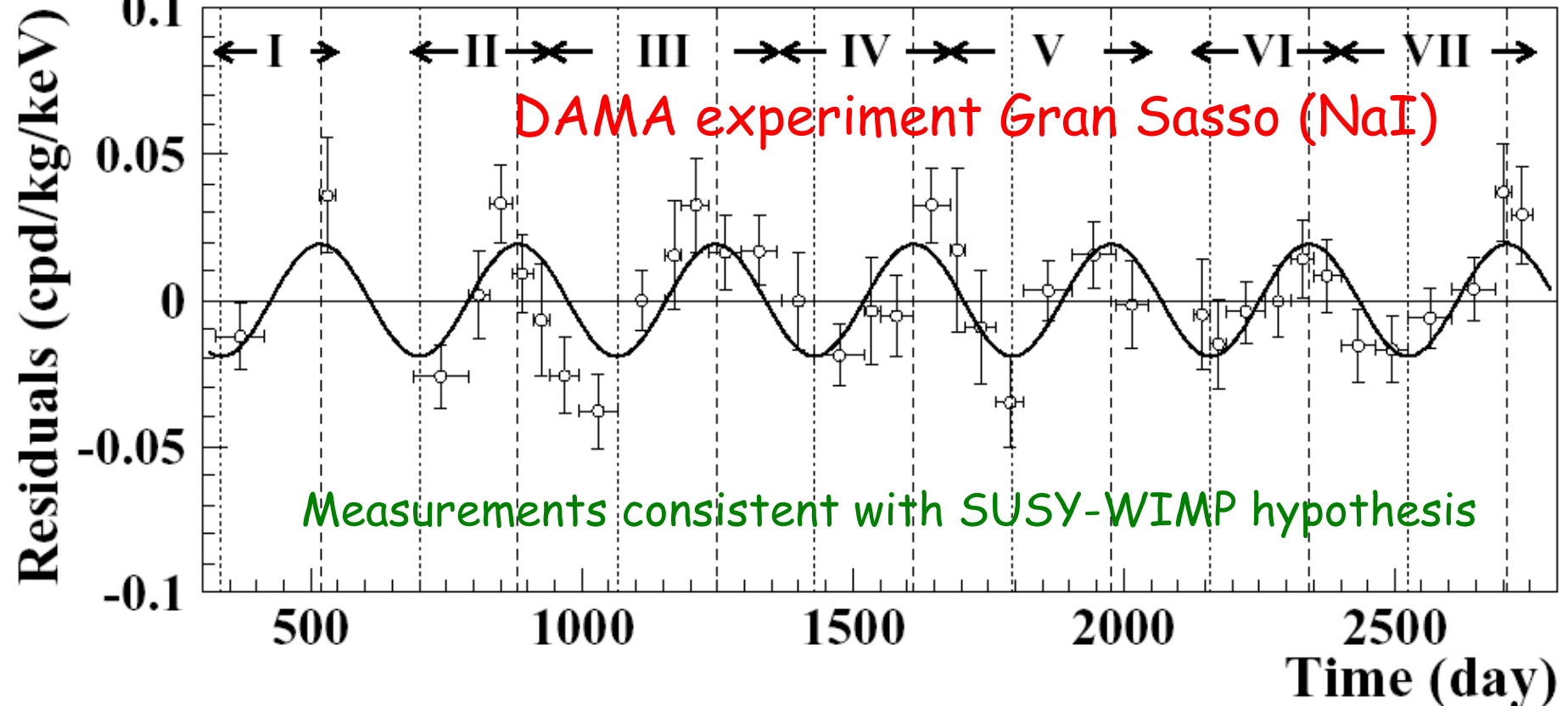
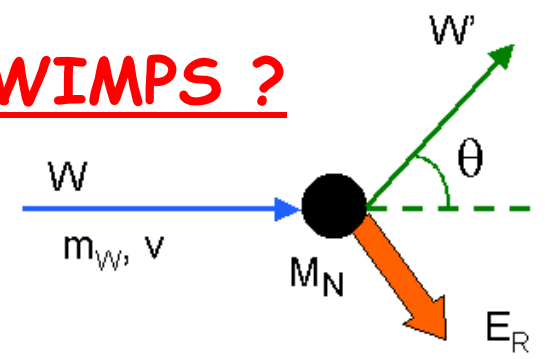
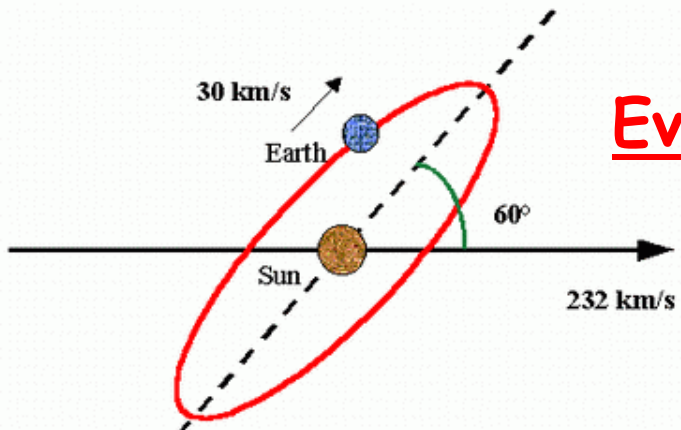
WMAP
measurement of
cosmic microwave
background

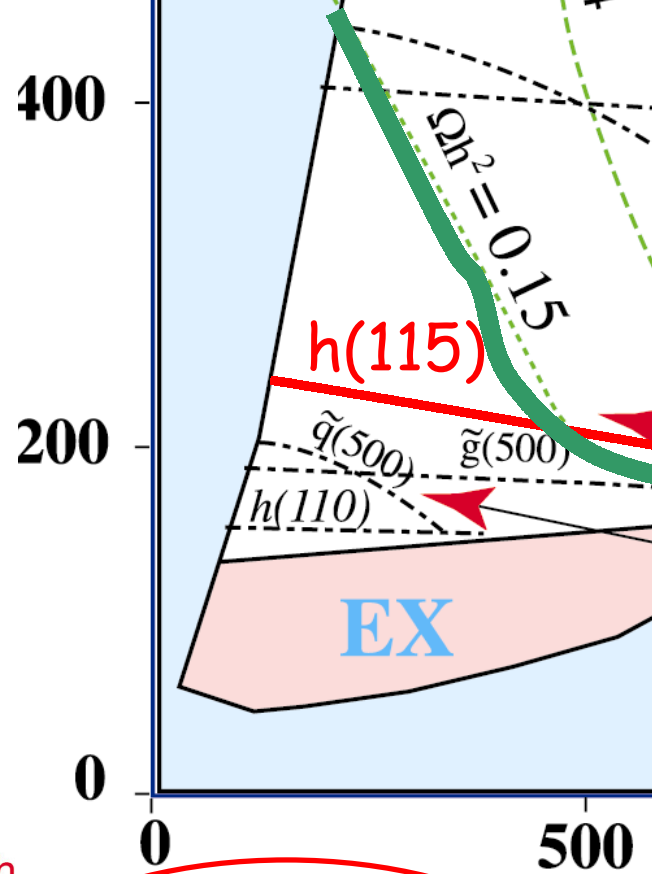
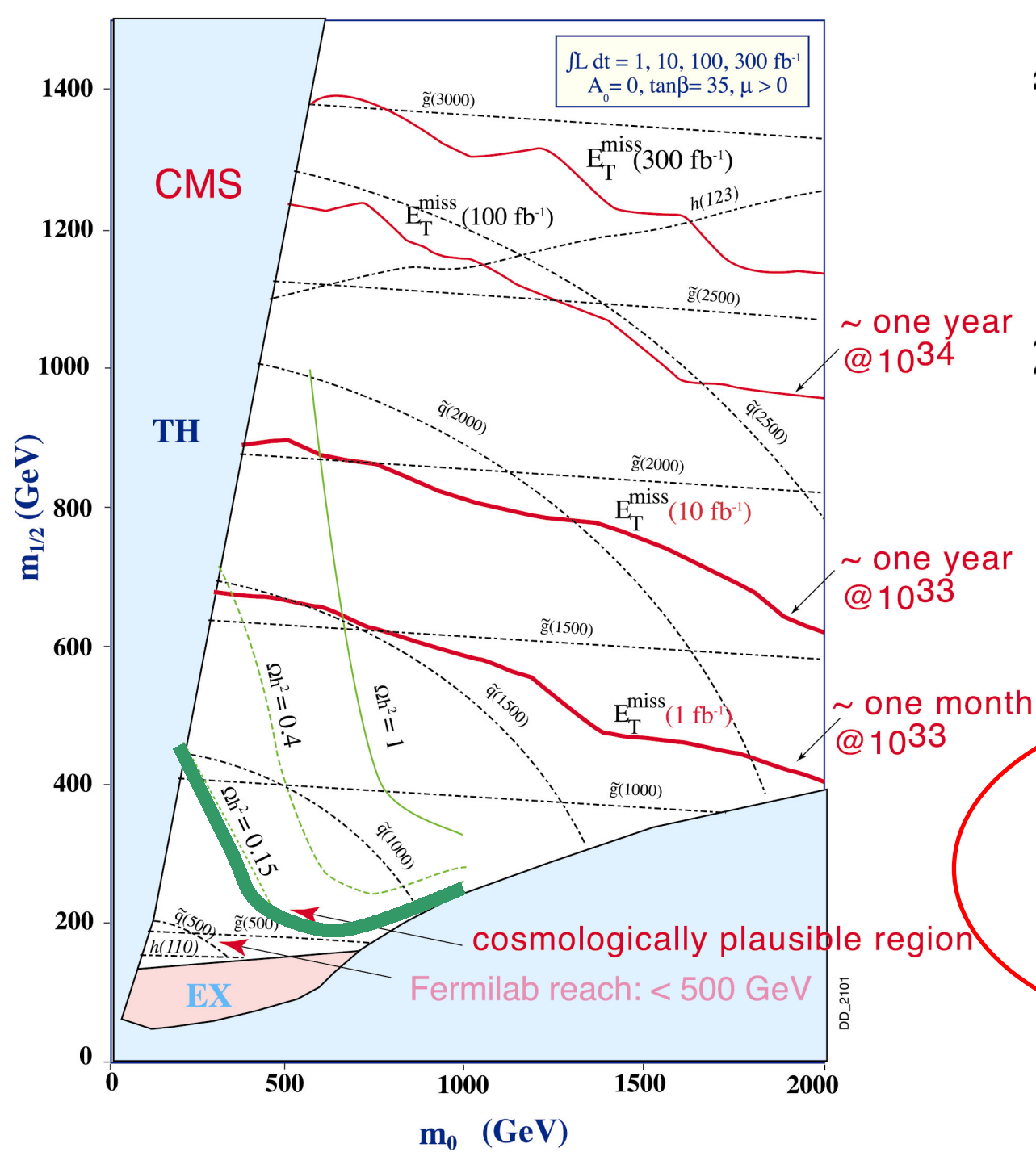


$$\Omega_{DM} = 0.23$$

Cosmological Constraints II

Evidence for relic WIMPS ?





If dark matter:
 Fermilab: **NO**
 LHC: **YES**

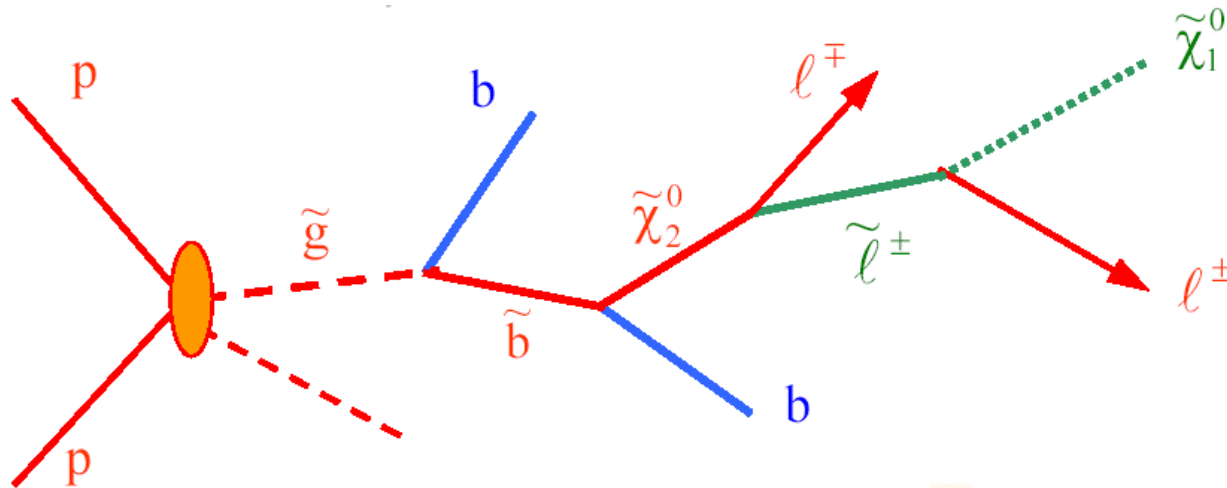
SUSY reach

SUSY search in pp

strategy:

high SM QCD background: jets

need something beyond: **leptons and/or missing energy**



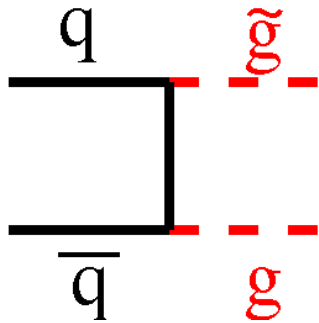
Examples:

- squarks and gluinos (missing energy) **strong**
- neutralinos and charginos (leptons and missing energy)

electroweak

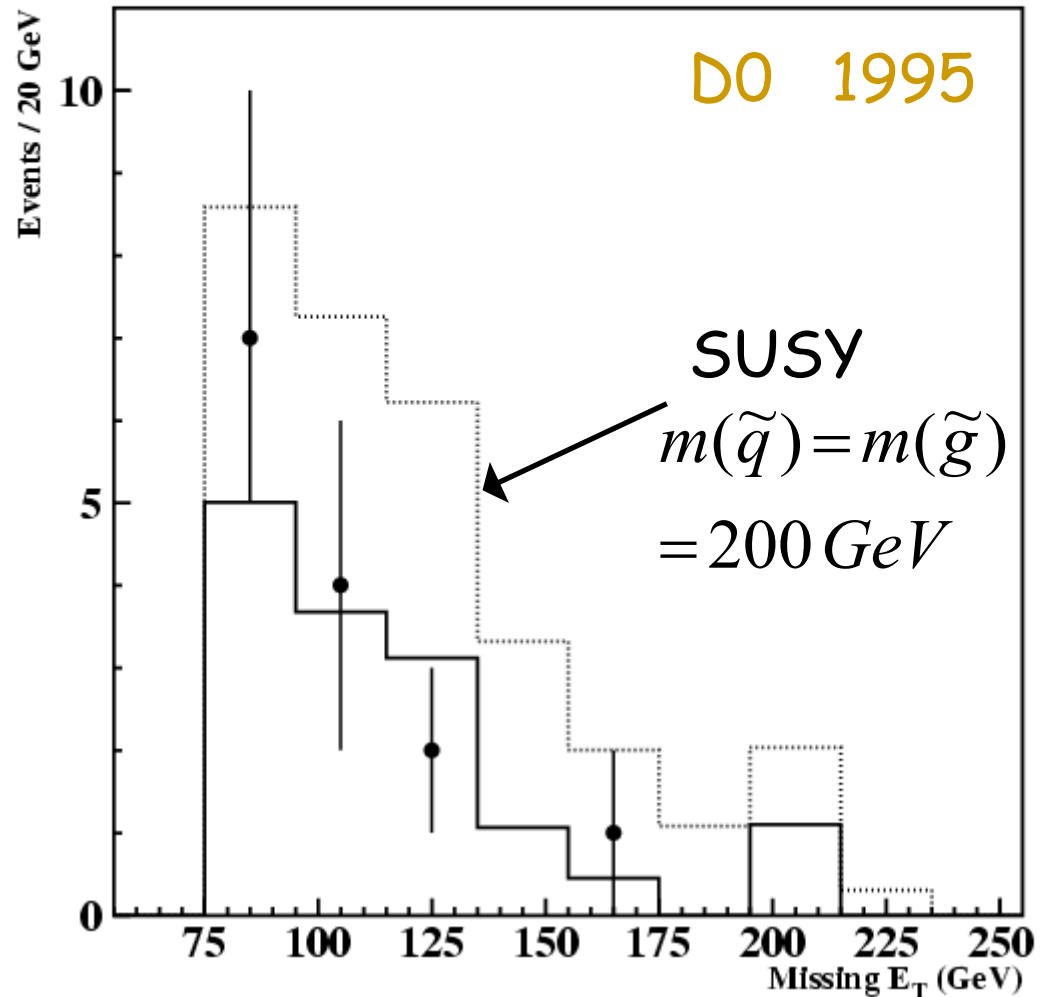
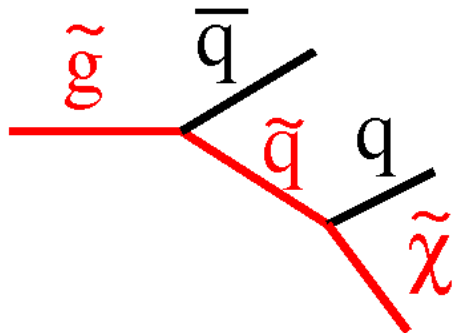
Jets + Missing Energy: Squark and Gluino Search

Production (ex.):

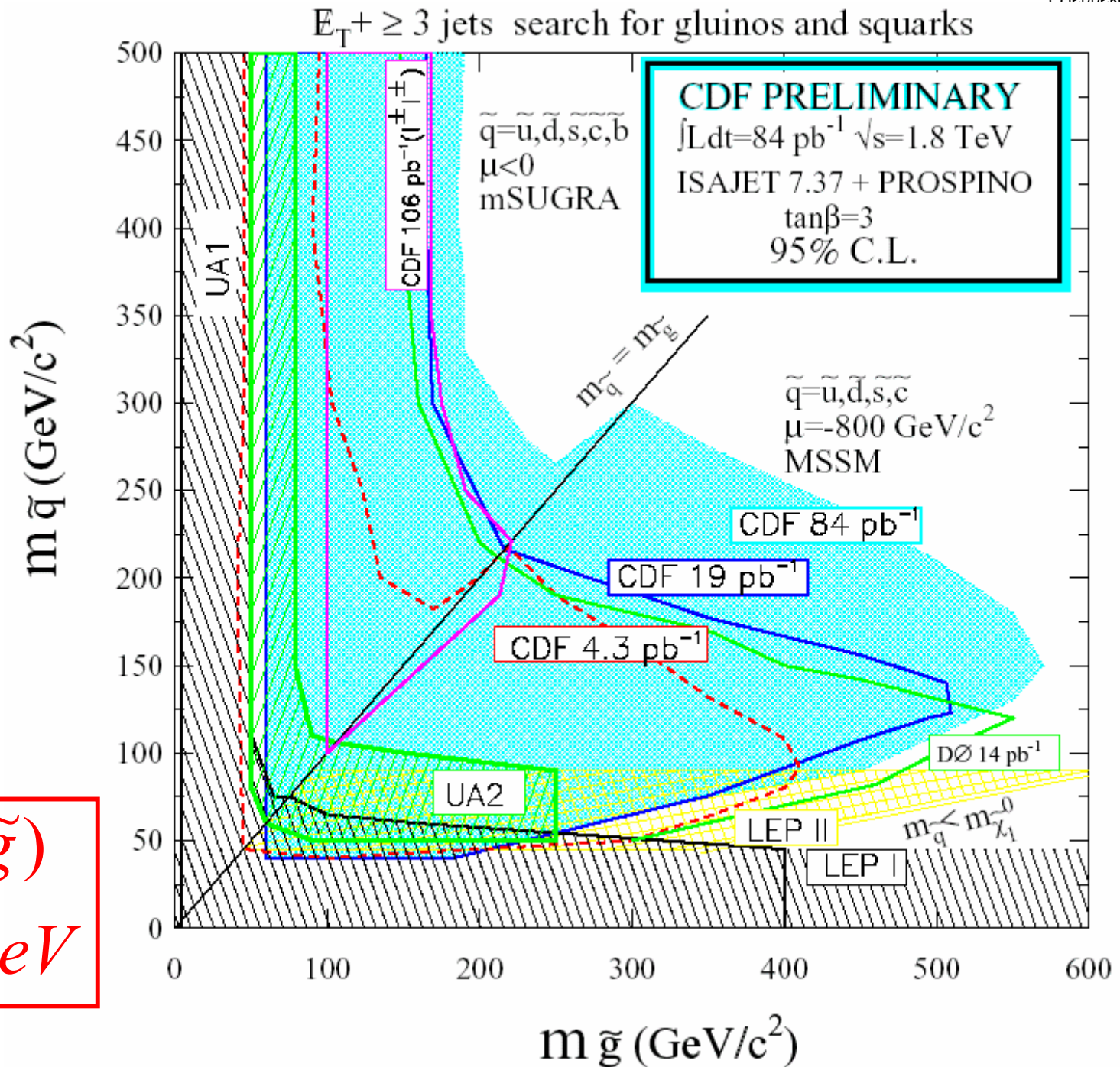


cross section large!

Decay (ex.):



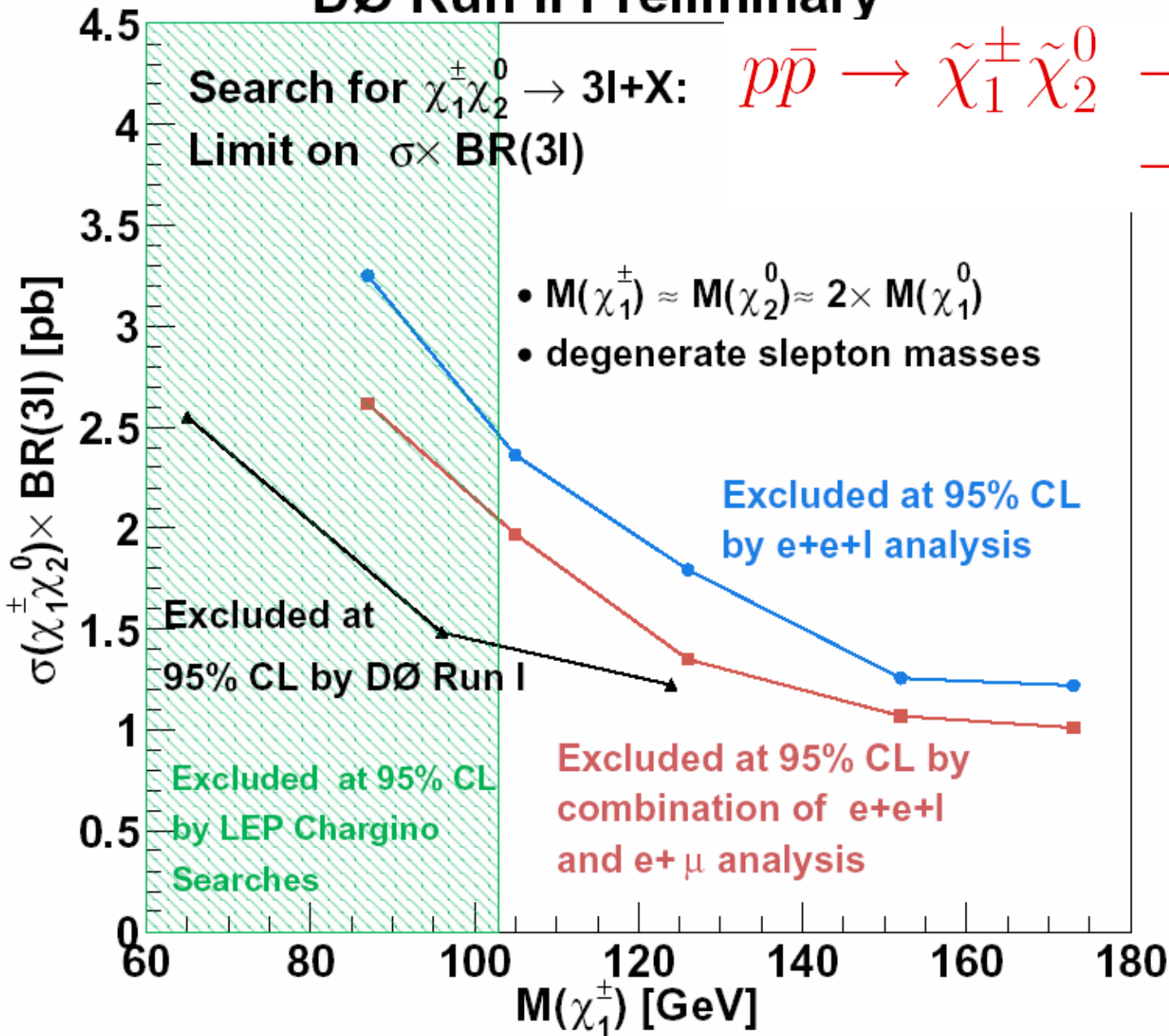
Results



$m(\tilde{q}), m(\tilde{g})$
 $> \sim 200 \text{ GeV}$

Leptons: Hunting Charginos and Neutralinos

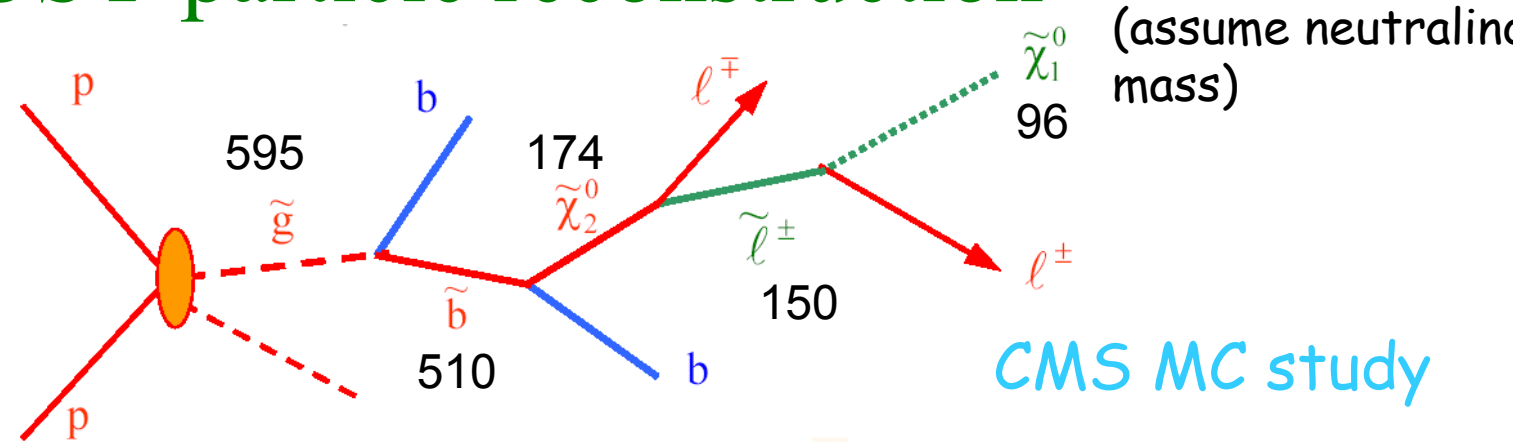
DØ Run II Preliminary



Expected
 SUSY
 xsection
 factor ~ 10
 lower

No
 (improved)
 SUSY
 limit yet

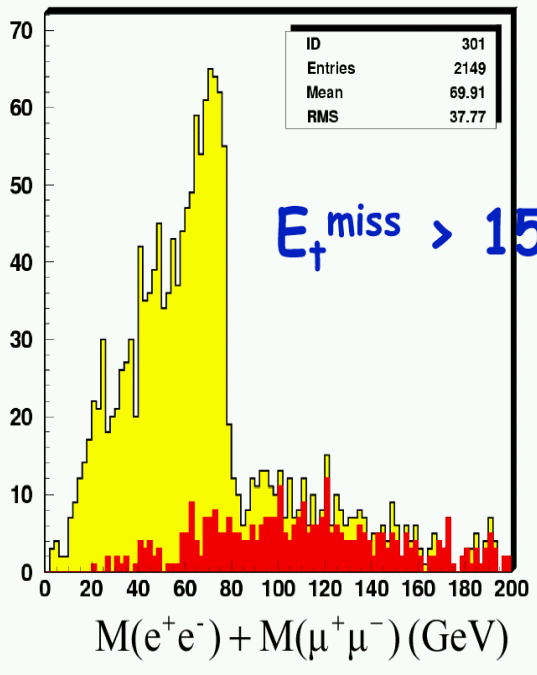
SUSY particle reconstruction



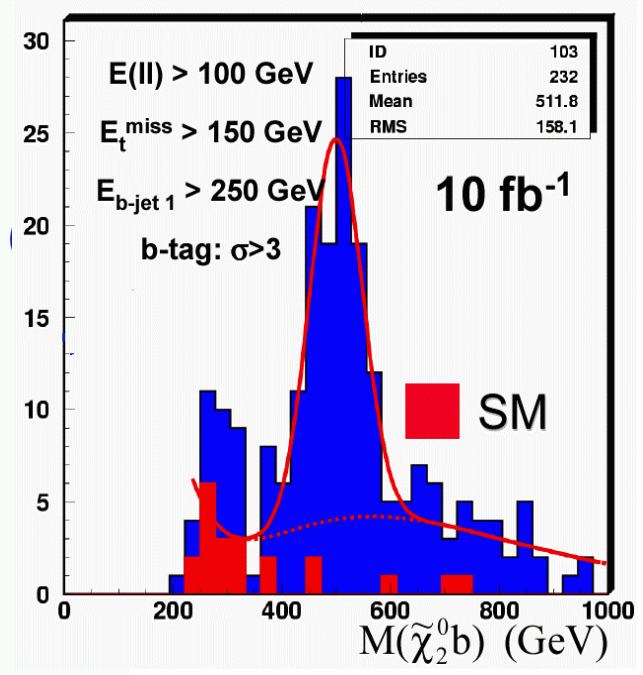
(neutralino 2)

sbottom

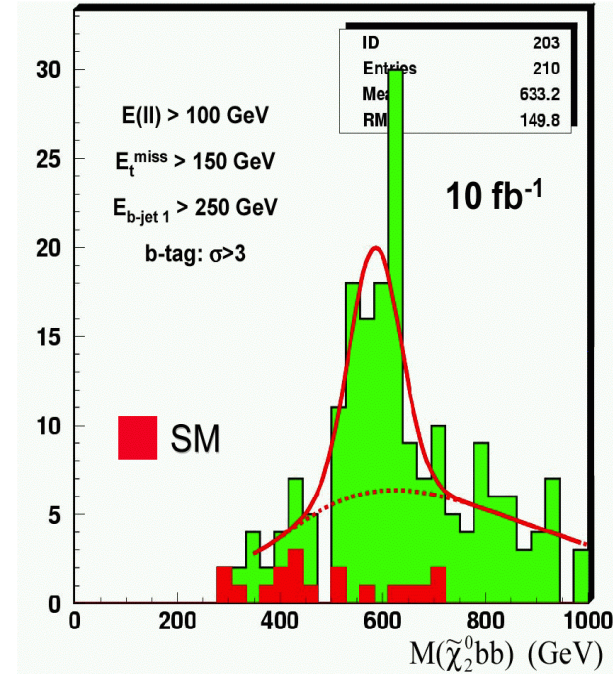
gluino



edge 78.9 +/- 2.1 GeV



499.4 +/- 6.6 GeV



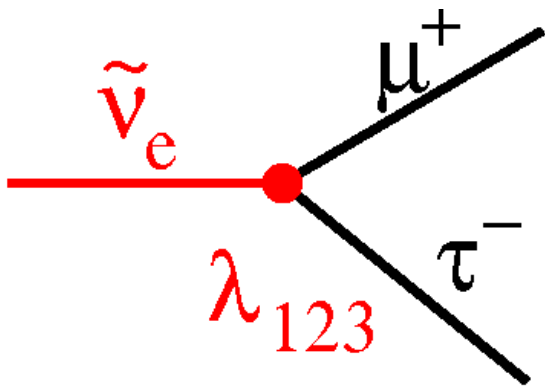
585.1 +/- 11.1 GeV

Chiorboli/Tricomi

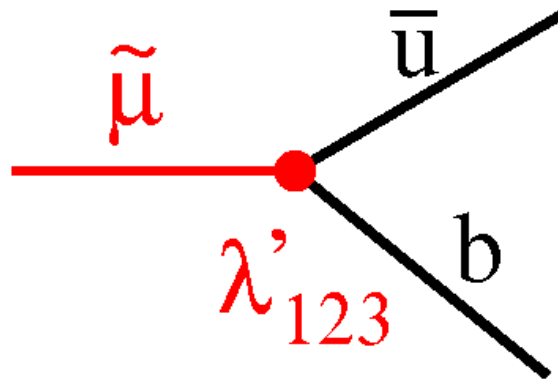
SUSY with R parity violation

- neutralino unstable, no dark matter candidate !
- lepton and/or baryon number violated

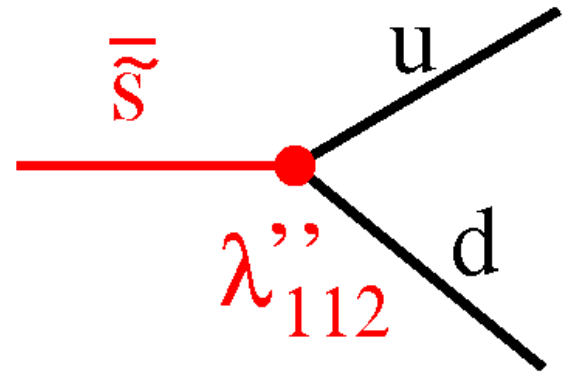
Example diagrams:



violates L



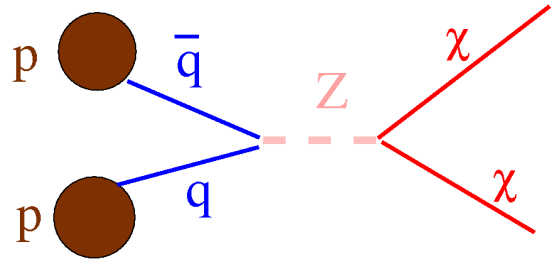
violates L



violates B

Example: Neutralino Decays via $\lambda_{121}, \lambda_{122}, \dots$

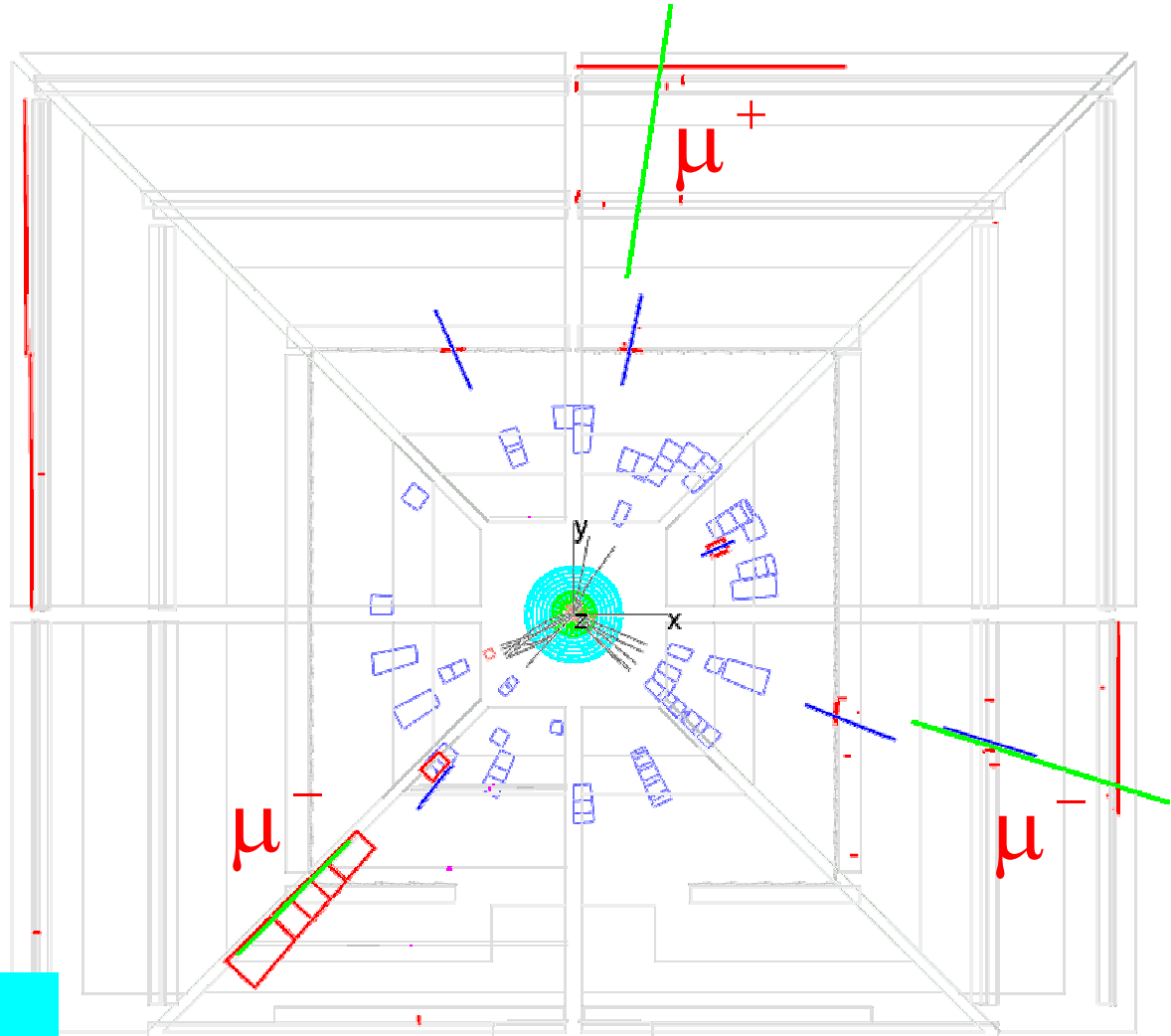
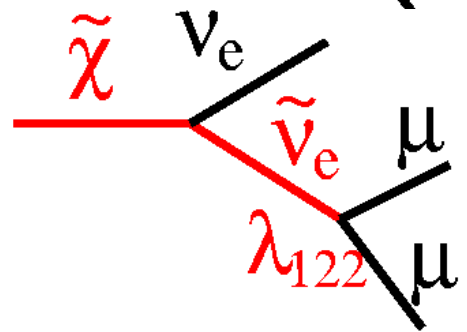
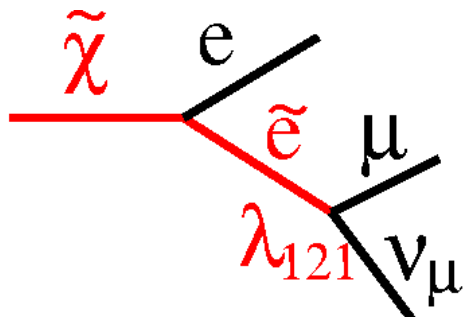
production:



Run 170246 Event 17918459 Tue Mar 4 18:23:36 2003

DO Run II

decay (examples):



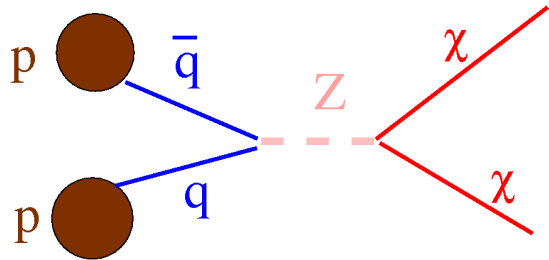
signature:

at least 3 charged leptons!

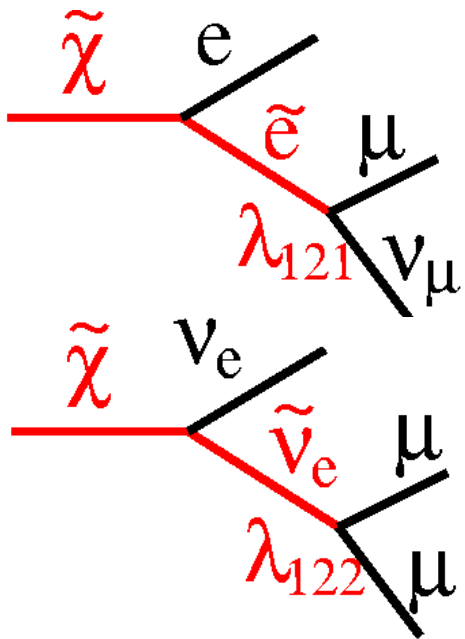
v 1, Front(X-Y)

Example: Neutralino Decays via $\lambda_{121}, \lambda_{122}, \dots$

production:



decay (examples):

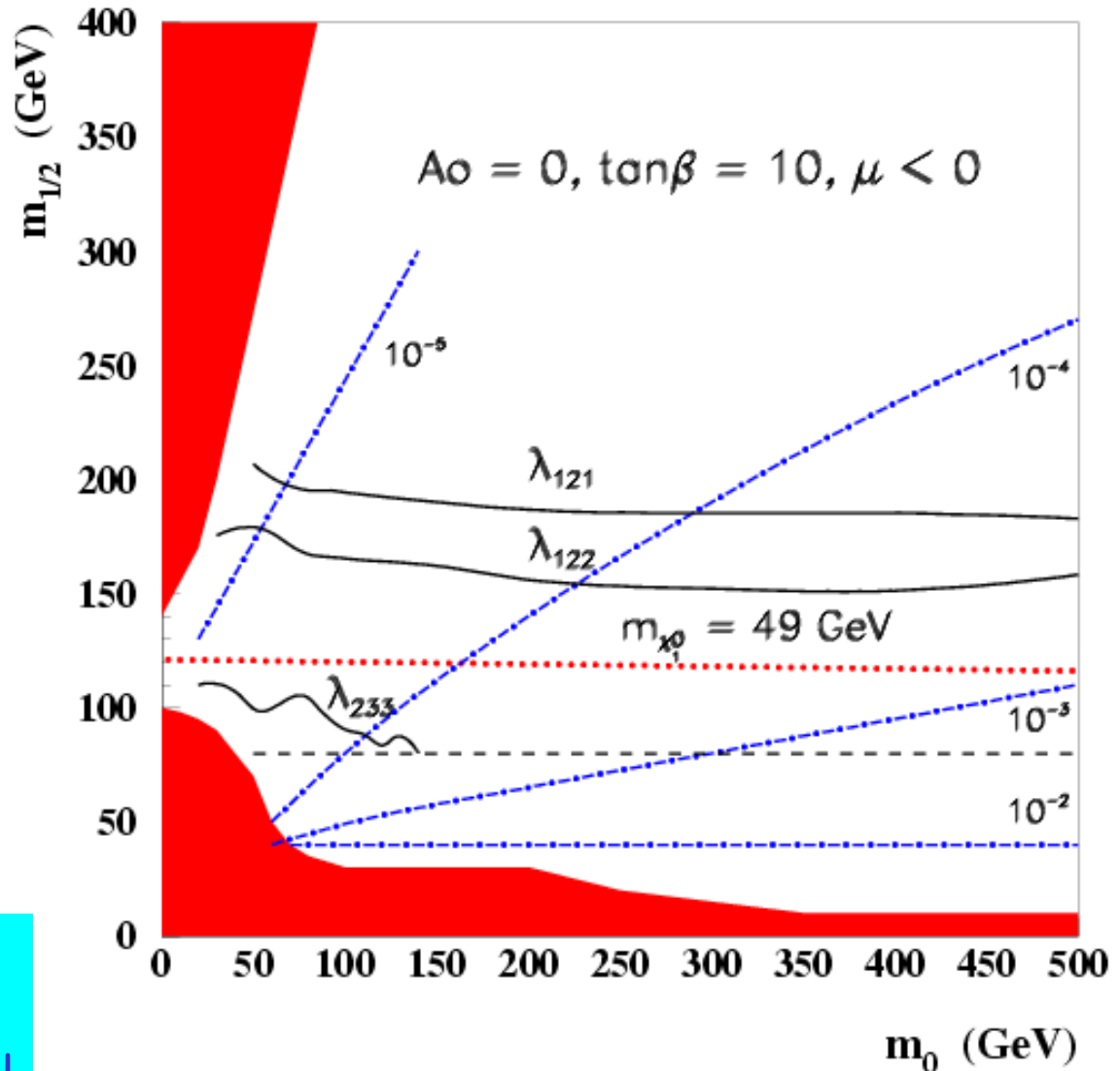


experimental signature:

at least 3 charged leptons!

upper limits:

D0 Run



(Large) Extra Dimensions

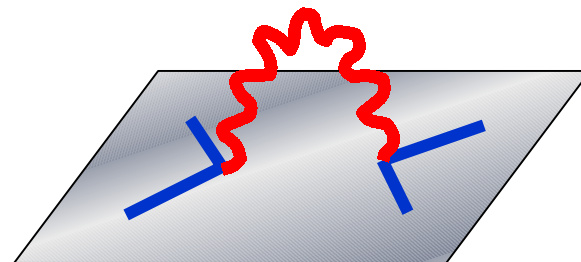
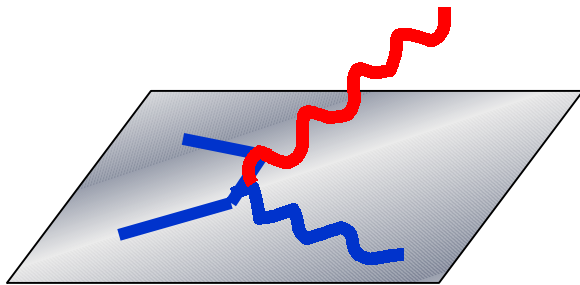
Why is **gravity** so different from the other interactions ?

mass and length scales:

$$M_{ew} \sim 10^2 \text{ GeV} \quad l_{ew} \sim 10^{-18} \text{ m}$$

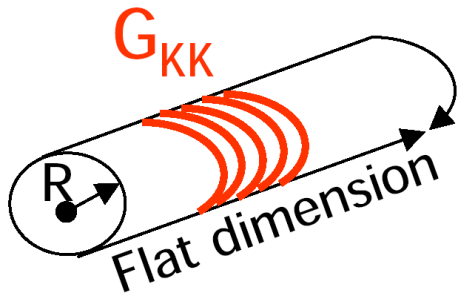
$$M_{Pl} \sim \frac{1}{\sqrt{G_N}} \sim 10^{19} \text{ GeV} \quad l_g \sim 10^{-35} \text{ m}$$

Idea: only one fundamental scale $M_S \sim 100 - 1000 \text{ GeV}$
 gravity appears weak since gravitons propagate
 in $4 + n$ dimensions („dilution“)



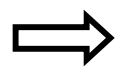
Extra Dimensions: phenomenology

n extra dimensions of space with size R:



Compactified dimension

$$V \sim \frac{1}{M_{Pl}^2} \frac{m_1 m_2}{r}$$



$$V \sim \frac{1}{M_S^{2+n}} \frac{1}{R^n} \frac{m_1 m_2}{r}$$



- deviations from Newton/Einstein laws for $r < R$

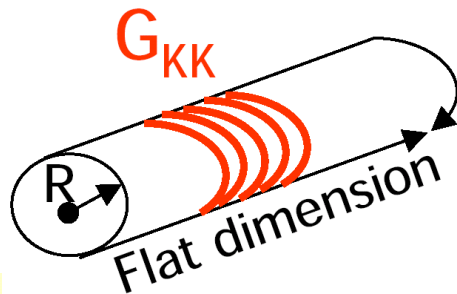
therefore n=1 and n=2 ruled out!

- gravitons G show up in high energy physics experiments as real or virtual particles

pp colliders!

Extra Dimensions: phenomenology

n extra dimensions of space with size R:



Compactified
dimension

$$R = \frac{1}{2\sqrt{\pi}M_S} \left(\frac{M_{Pl}}{M_S} \right)^{2/n}$$

$$8 \times 10^{12} \text{ m}, \quad n = 1$$

$$0.7 \text{ mm}, \quad n = 2$$

$$3 \text{ nm}, \quad n = 3$$

$$6 \times 10^{-12} \text{ m}, \quad n = 4$$



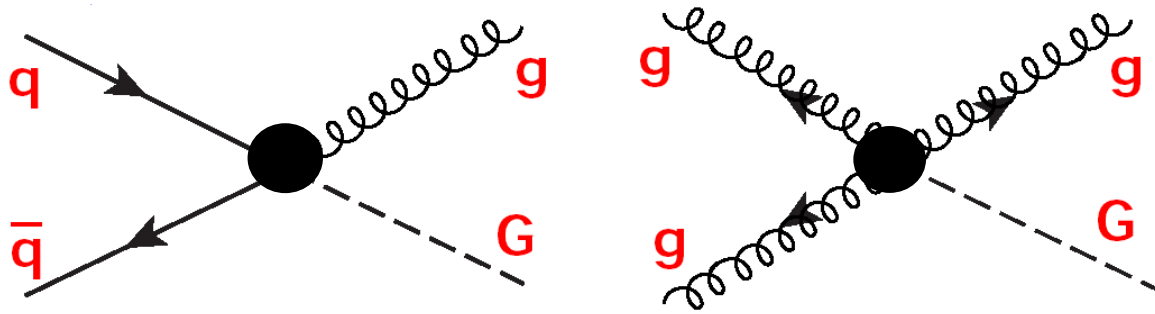
Escher

- deviations from Newton/Einstein laws for $r < R$

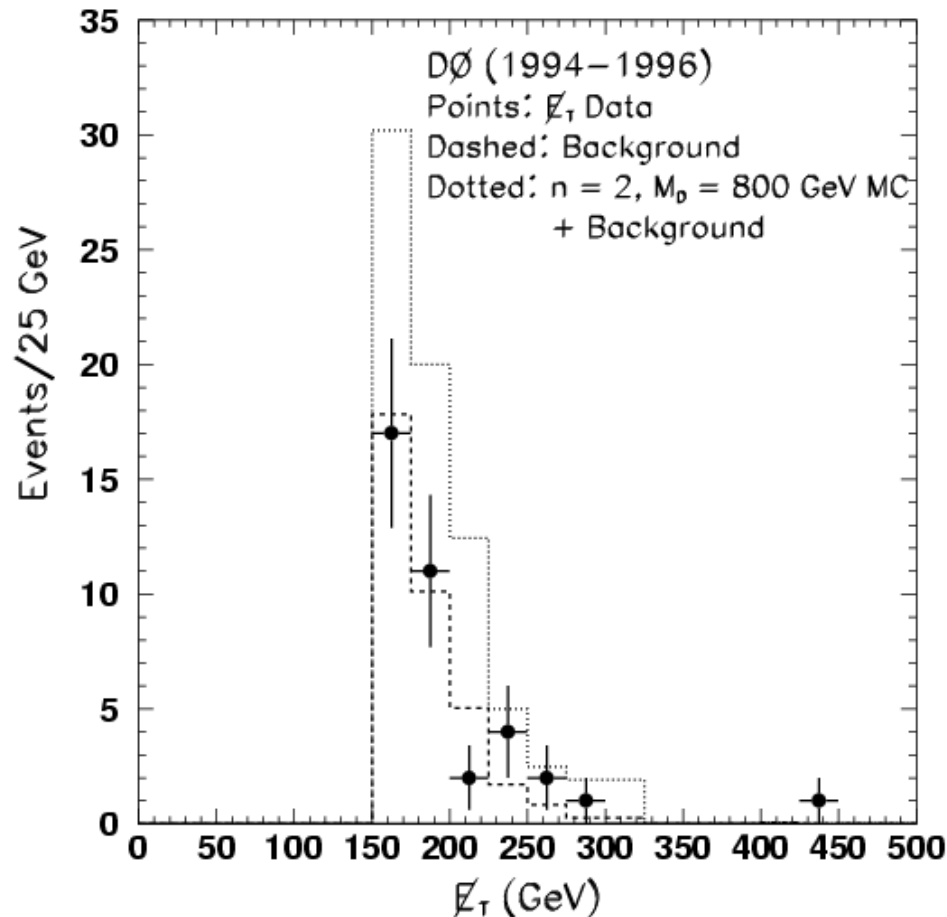
therefore n=1 and n=2 ruled out!

- gravitons G show up in high energy physics experiments as real or virtual particles } **pp colliders!**

Real Graviton Emission in $p p$



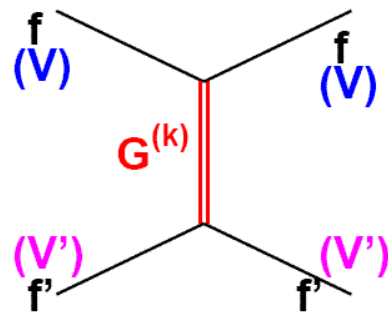
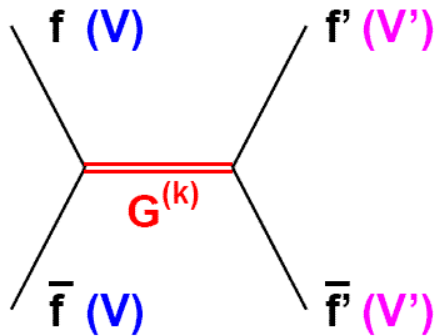
Monojet signatures!



resulting lower limits on M_S/GeV :

$n=2$	$n=3$	$n=4$	$n=5$	$n=7$
890	730	680	640	620

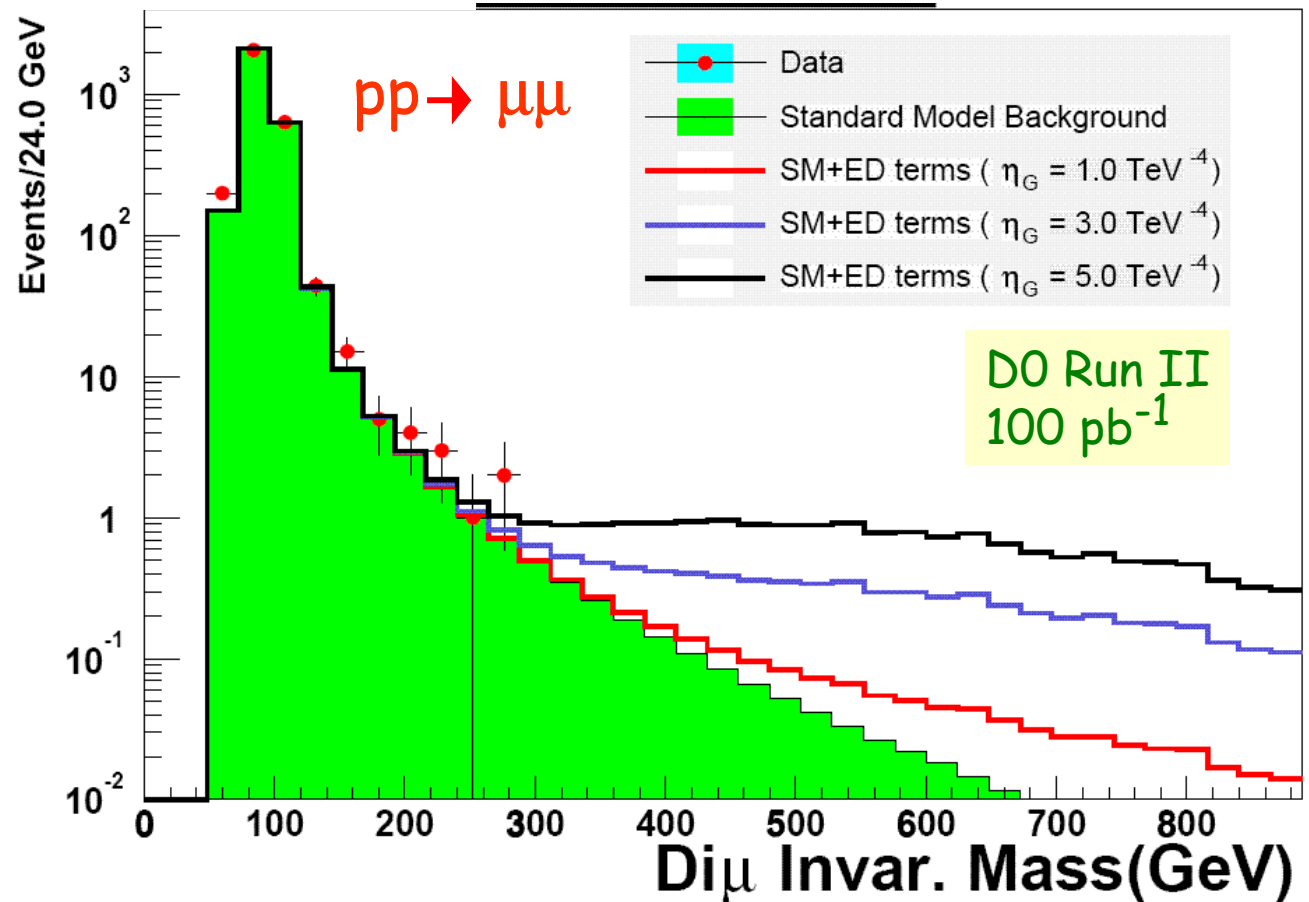
Virtual Graviton effects in $p p$



SM cross sections modified!

All D0 analyses
($ee, \mu\mu, \gamma\gamma$)
combined:

$$M_S(n=2) > 1.38 \text{ TeV}$$



Black holes ?

predicted in large extra dimension models

production: mass 1 - 10 TeV, xsection large (\sim nb)

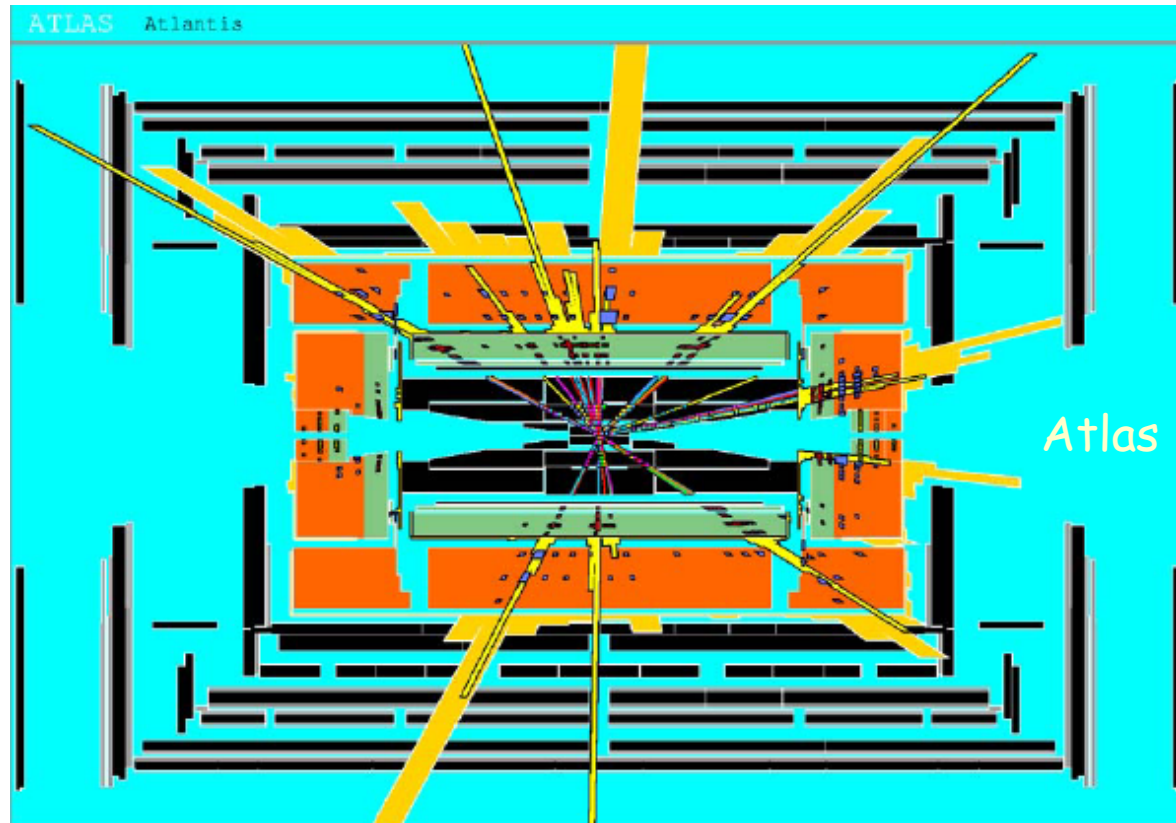
decay:

Hawking radiation

All SM d.o.f. equally likely

Multiplicity up to 30

Decay also into higgs!



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Part IV New Phenomena

- SUSY
 - motivation
 - searches:
 - R-Parity conserved
 - R-Parity violated
- Extra dimensions
- Black holes

References

References

- lectures:

F. Gianotti, LHC physics

www.wlap.org/cern/lectures/summer/2000/gianotti

J. Womersley, Physics at Hadron Colliders

d0server1.fnal.gov/users/womersley/brazil1.pdf...brazil4.pdf

- experimental homepages:

www-cdf.fnal.gov

www-d0.fnal.gov

atlas.web.cern.ch

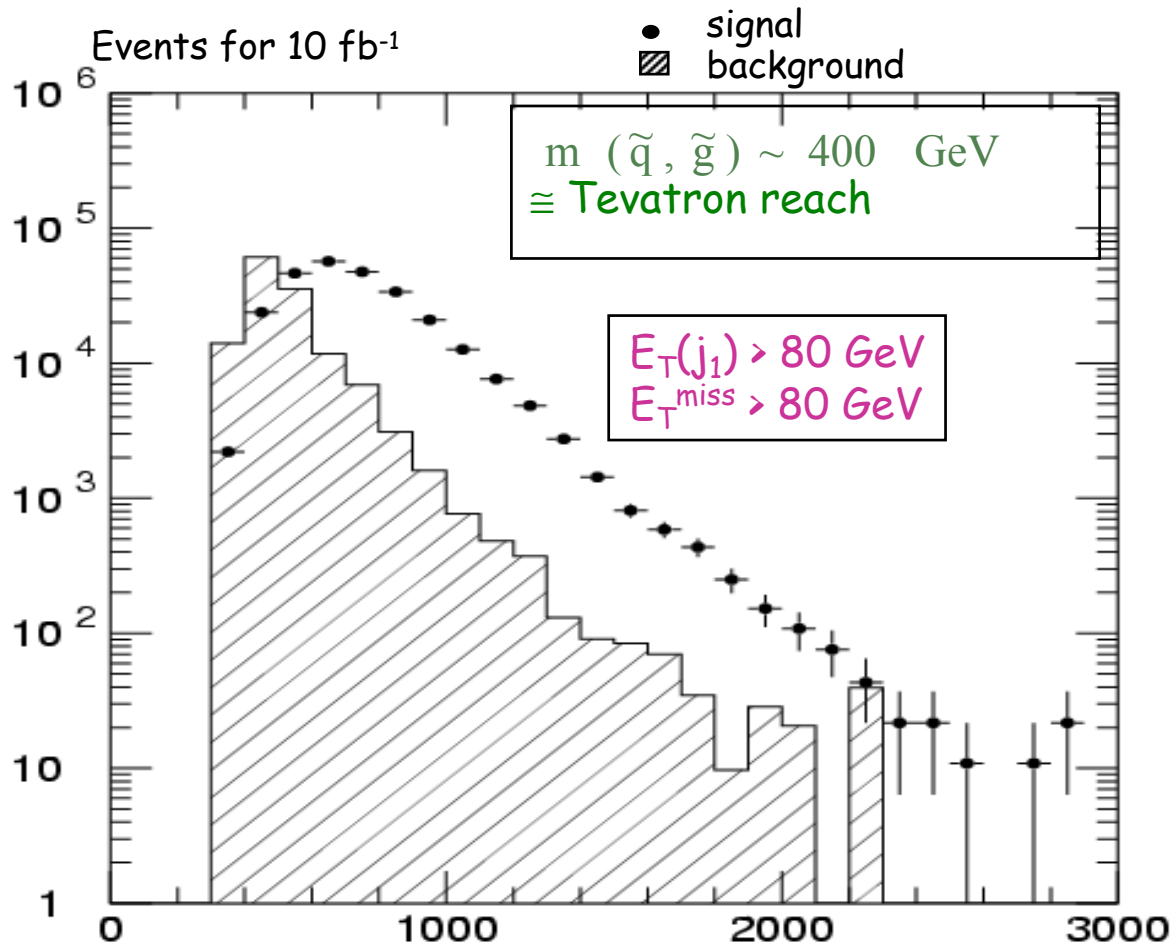
cmsinfo.cern.ch

- theory:

Physics at Run II workshop

fnth37.fnal.gov/run2.html

Appendices



$$M_{\text{eff}} = E_T^{\text{miss}} + \sum_{i=1}^4 p_T(\text{jet}_i) \quad (\text{GeV})$$

Inclusive $\ell^+\ell^- + E_t^{\text{miss}}$ final states

$m_0 = 200 \text{ GeV}$, $m_{1/2} = 160 \text{ GeV}$,
 $\text{tg}\beta = 2$, $A_0 = 0$, $\mu < 0$

